

## Spectroscopic data on the $^{25}\text{Mg} \rightarrow ^{24}\text{Mg} + n$ excited configurations from the $^{24}\text{Mg}(d,p)^{25}\text{Mg}$ reaction.

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The aim of this work was to obtain precise experimental differential cross sections (DS) of the neutron transfer reaction  $^{24}\text{Mg}(d,p)^{25}\text{Mg}$  for subsequent extraction of asymptotic normalization coefficients (ANC) and spectroscopic factors (SF). These data are necessary both to clarify the structure of light nuclei and to calculate the astrophysical S-factors of processes in the MgAl cycle of stellar proton burning:  $^{24}\text{Mg}(p,\gamma)^{25}\text{Al}(\beta+)^{25}\text{Mg}(p,\gamma)^{26}\text{Al}(\beta+)^{26}\text{Mg}$ , using the relationship between the mirror states of  $^{25}\text{Mg}$  and  $^{25}\text{Al}$  nuclei.

The DS of the  $^{24}\text{Mg}(d,p)^{25}\text{Mg}$  reaction have been measured in the forward hemisphere of the angles for the neutron stripping process to the states  $E^* = 0.585 \text{ MeV}, 1/2^+; 0.975 \text{ MeV}, 3/2^+; 1.61 \text{ MeV}, 7/2^+; 1.96 \text{ MeV}, 5/2^+$  and  $2.56 \text{ MeV}, 1/2^+$  of the  $^{25}\text{Mg}$  nucleus at  $E_d = 14.5 \text{ MeV}$ . The DS for the  $E^* = 0.0 \text{ MeV}, 5/2^+$  is given in work [1] as well as the description of experiment which was carried out at the deuteron beam of the U-150M accelerator of INP RKaz.

The experimental DS have been analyzed in the framework of the modified distorted wave Born approximation (MDWBA) [2] using the DWUCK5 [3] code. Our previous analysis [4] has shown that the neutron transfer to the ground state of the  $^{25}\text{Mg}$  nucleus is non-peripheral process. Here it is shown that this process becomes more peripheral with increasing the excitation energy of the final nucleus, which allowed us to estimate the values of the neutron coupling ANC for the above-mentioned excited states of the  $^{25}\text{Mg}$  nucleus and evaluate the dependence of extracted SF values on the geometry parameters of the Woods-Saxon neutron binding potential.

The obtained results will be further used to calculate the astrophysical S-factors of the radiative capture process  $^{24}\text{Mg}(n,\gamma)^{25}\text{Mg}$  at astrophysical relevant energies.

### References

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