

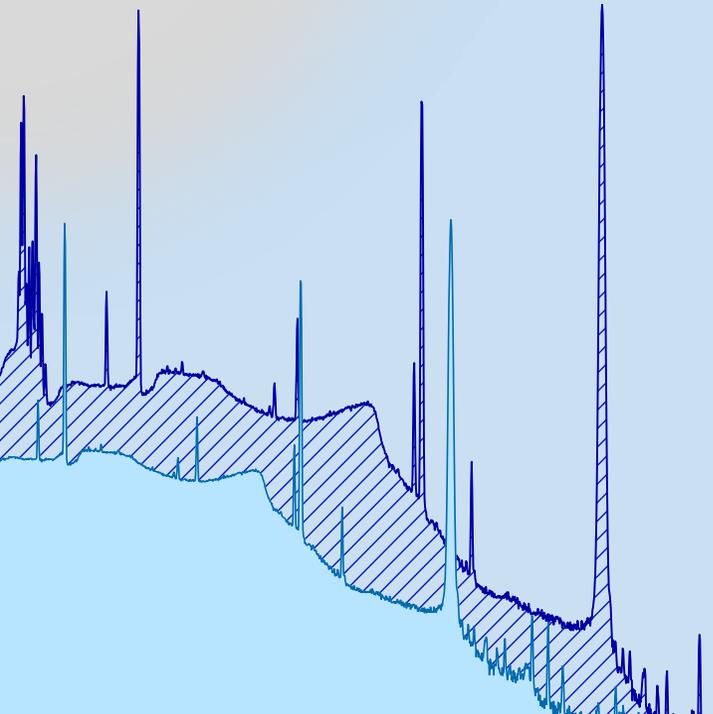


LXXV Международная конференция «ЯДРО-2025. Физика атомного
ядра и элементарных частиц. Ядерно-физические технологии»



«Excitation of isomeric states of Hg and Au isotopes in photonuclear reactions»

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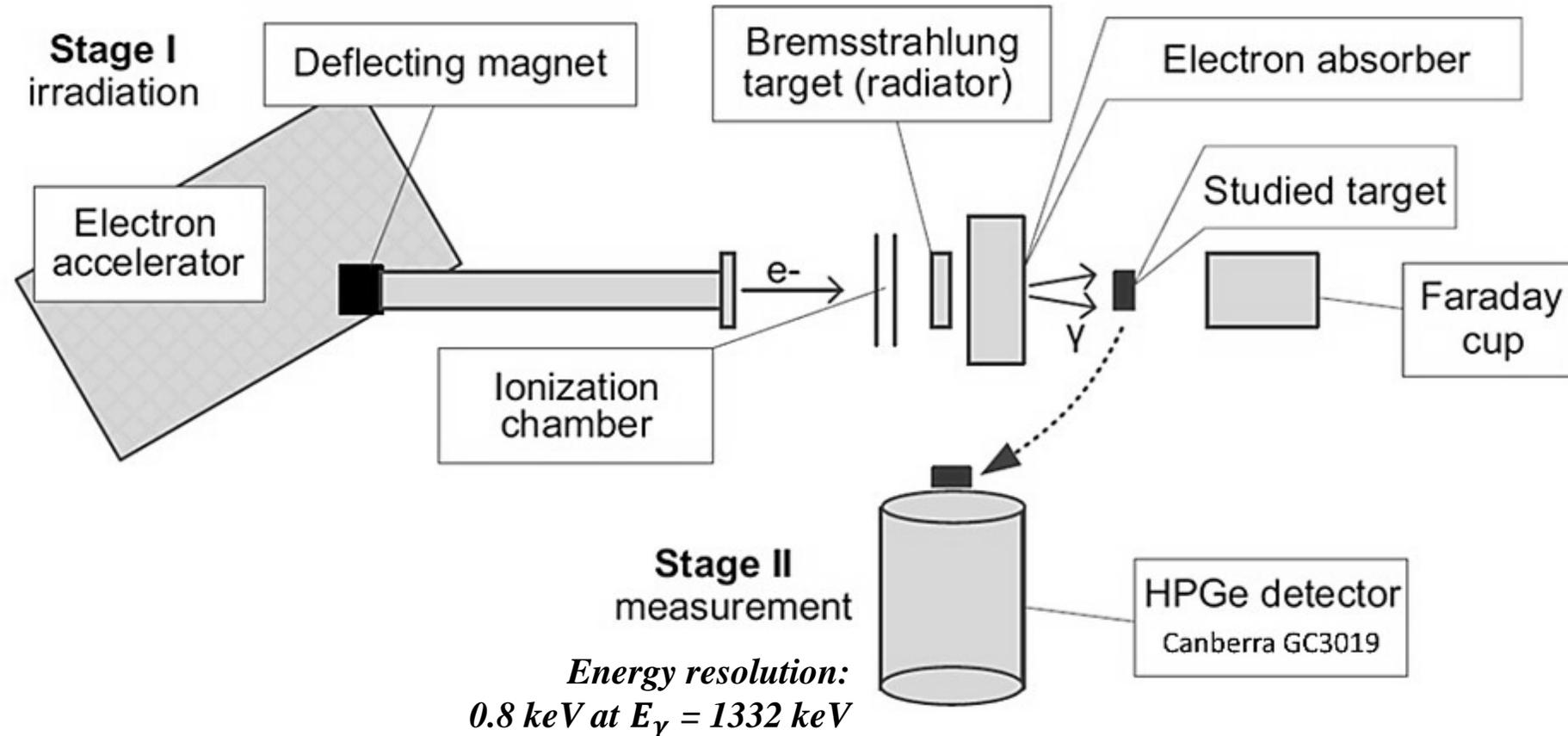
Saint Petersburg - 2025

The irradiation method of **a natural mixture of Hg isotopes** sample with bremsstrahlung radiation with $E^m = 55 \text{ MeV}$

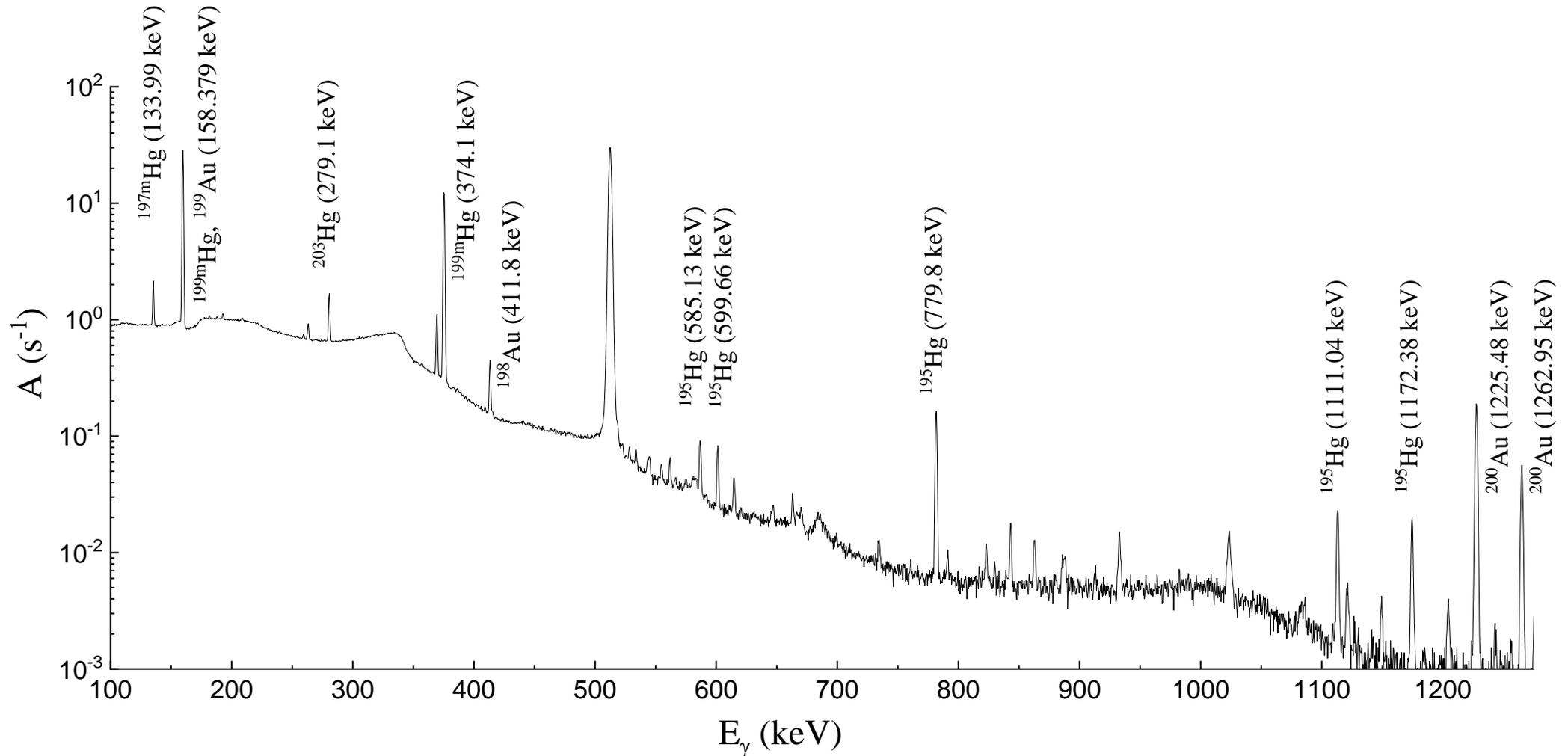
Converter target thickness: 0.2 mm

Irradiation time: 670 seconds

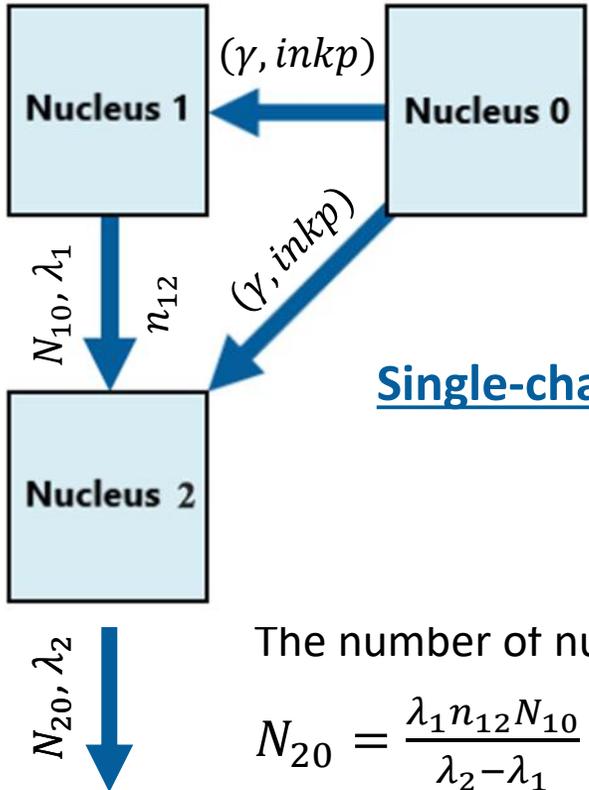
Thickness of the irradiated target: 2.8 mm



The yields of photonuclear reactions were determined from **the peaks of the γ -lines in the spectra of the residual activity** of the irradiated mercury sample



The spectrum of residual activity of the irradiated Hg sample (data set was carried out for 4 hours)

Determination of the yields of **single-channel** and **two-channel** photonuclear reactions


The number of nuclei at the end of irradiation: $N_{10} = \frac{S}{kI_{\gamma}(e^{-\lambda_1(t_2-t_1)} - e^{-\lambda_1(t_3-t_1)})}$

I_{γ} – intensity of the γ -line
 k – detector efficiency
 S – the peak area in the spectrum of residual activity during the measurement
 I – accelerator current

Single-channel reaction yield:

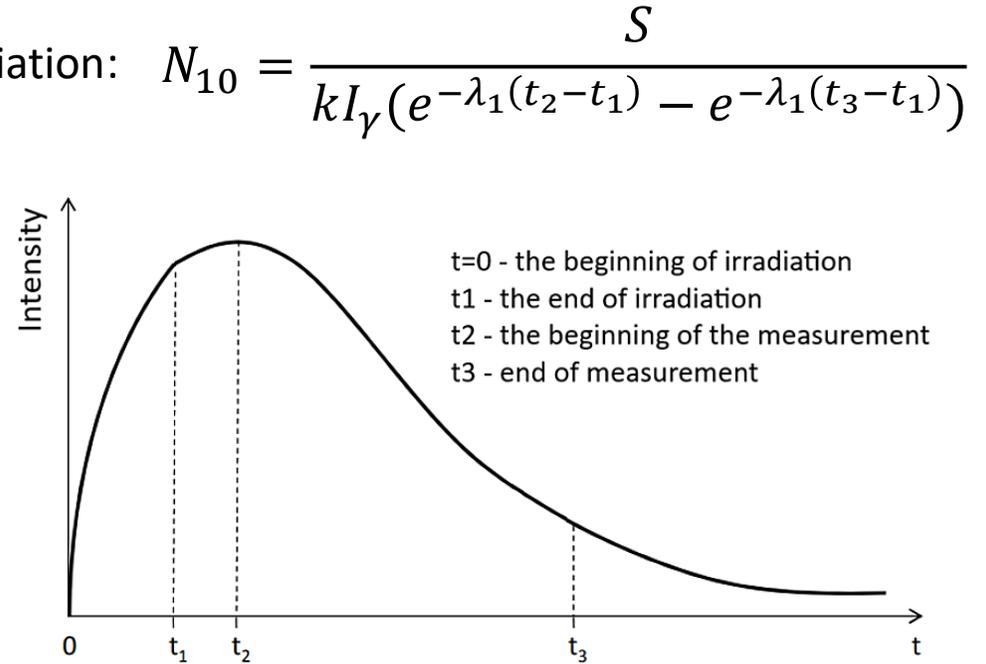
$$Y_1^{exp} = \frac{N_{10}\lambda_1}{I(1 - e^{-\lambda_1 t_1})}$$

The number of nuclei at the end of irradiation:

$$N_{20} = \frac{\lambda_1 n_{12} N_{10}}{\lambda_2 - \lambda_1} + \frac{S}{kI_{\gamma}(e^{-\lambda_2(t_2-t_1)} - e^{-\lambda_2(t_3-t_1)})} - \frac{\lambda_2 n_{12} N_{10}}{\lambda_2 - \lambda_1} \left(\frac{e^{-\lambda_1(t_2-t_1)} - e^{-\lambda_1(t_3-t_1)}}{e^{-\lambda_2(t_2-t_1)} - e^{-\lambda_2(t_3-t_1)}} \right),$$

Two-channel reaction yield:

$$Y_2^{exp} = \frac{\lambda_2}{I(1 - e^{-\lambda_2 t_1})} N_{20} - n_{12} Y_1 \left(\frac{\lambda_2(1 - e^{-\lambda_1 t_1}) - \lambda_1(1 - e^{-\lambda_2 t_1})}{(\lambda_2 - \lambda_1)(1 - e^{-\lambda_2 t_1})} \right), \text{ where } n_{12} \text{ – the decay coefficient}$$



Based on the obtained **yields, the isomeric ratios IR** of photonuclear reactions were calculated for the studied isotopes of $^{195,197}\text{Hg}$ and $^{198,200}\text{Au}$ at $E^m = 55 \text{ MeV}$

$$\text{IR} = Y_h/Y_l \quad \text{or} \quad \text{IR} = \sigma_h/\sigma_l,$$

where Y_h , Y_l and σ_h , σ_l – yields and cross sections of the formation of high-spin and low-spin states, respectively

The isomerism of atomic nuclei is caused by a **large difference in the spins** or **deformations of the isomeric state relative to the ground state.**

The population probability of isomers in photonuclear reactions **depends on:**

1. The reaction energy
2. The orbital angular momentum of emitted particles (which depends on their energy and the shell structure of the excited nucleus)
3. The spin-parity characteristics of the final state
4. Probabilities of cascading transitions from higher-lying states

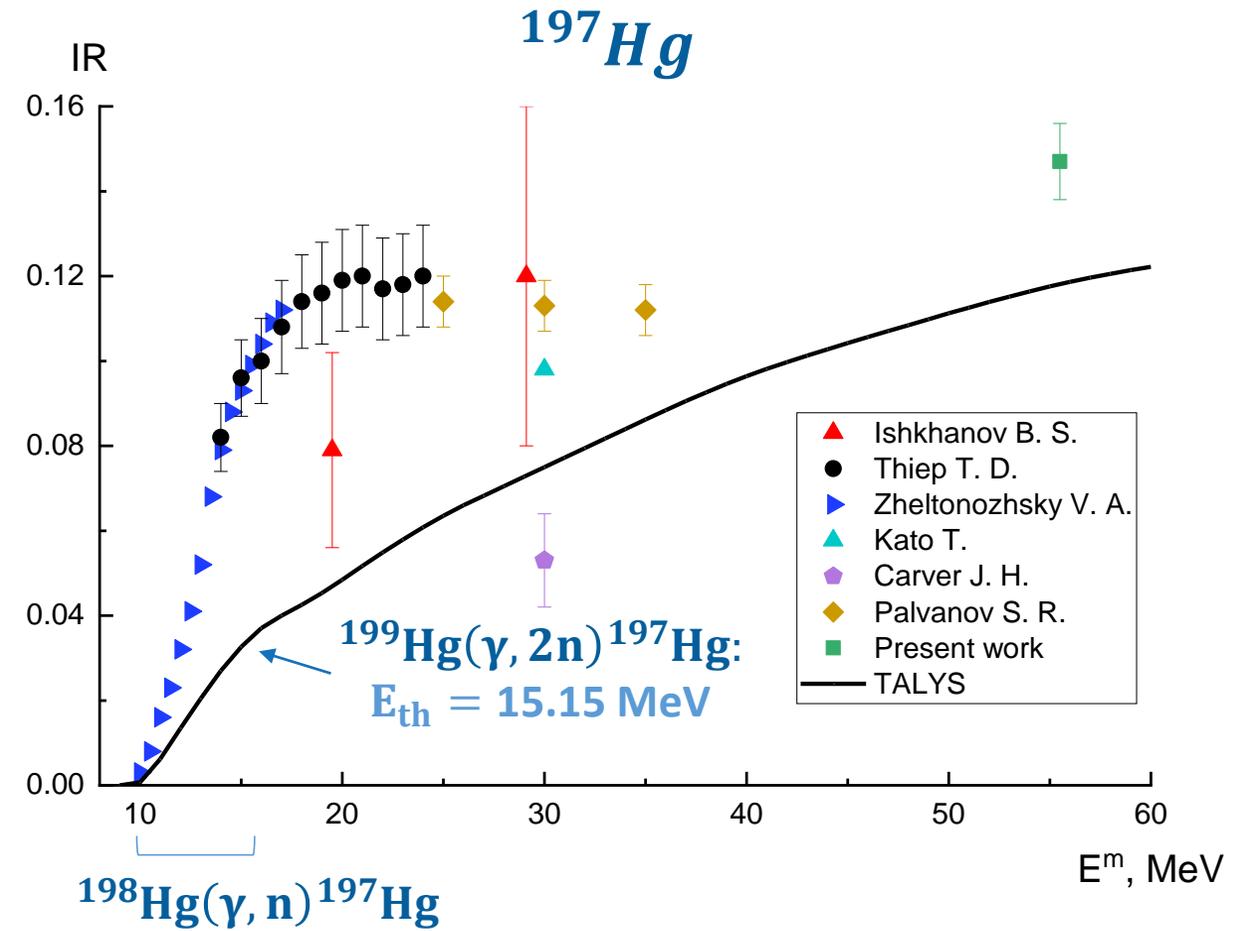
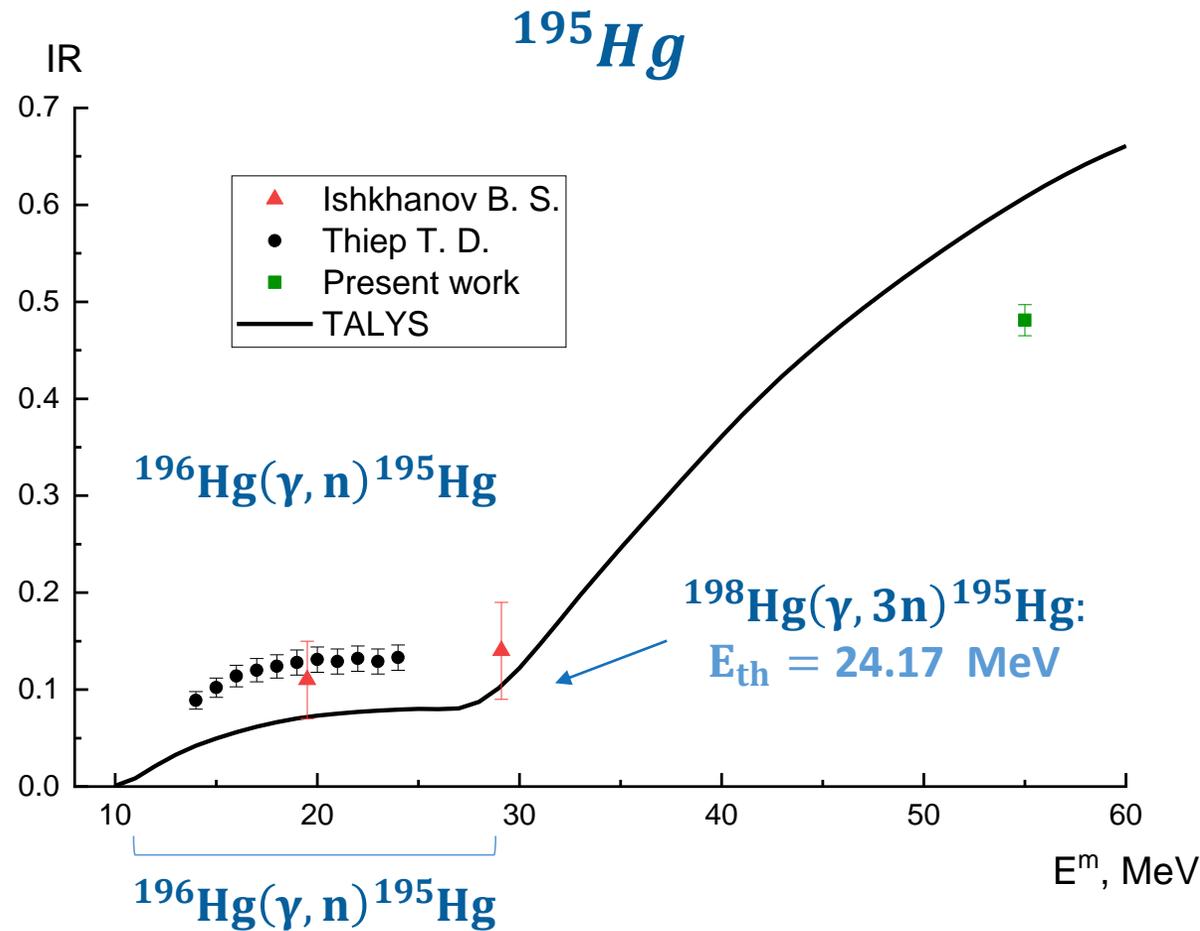
Based on the obtained yields, the isomeric ratios IR of photonuclear reactions were calculated for the studied isotopes of $^{195,197}\text{Hg}$ and $^{198,200}\text{Au}$ at $E^m = 55 \text{ MeV}$

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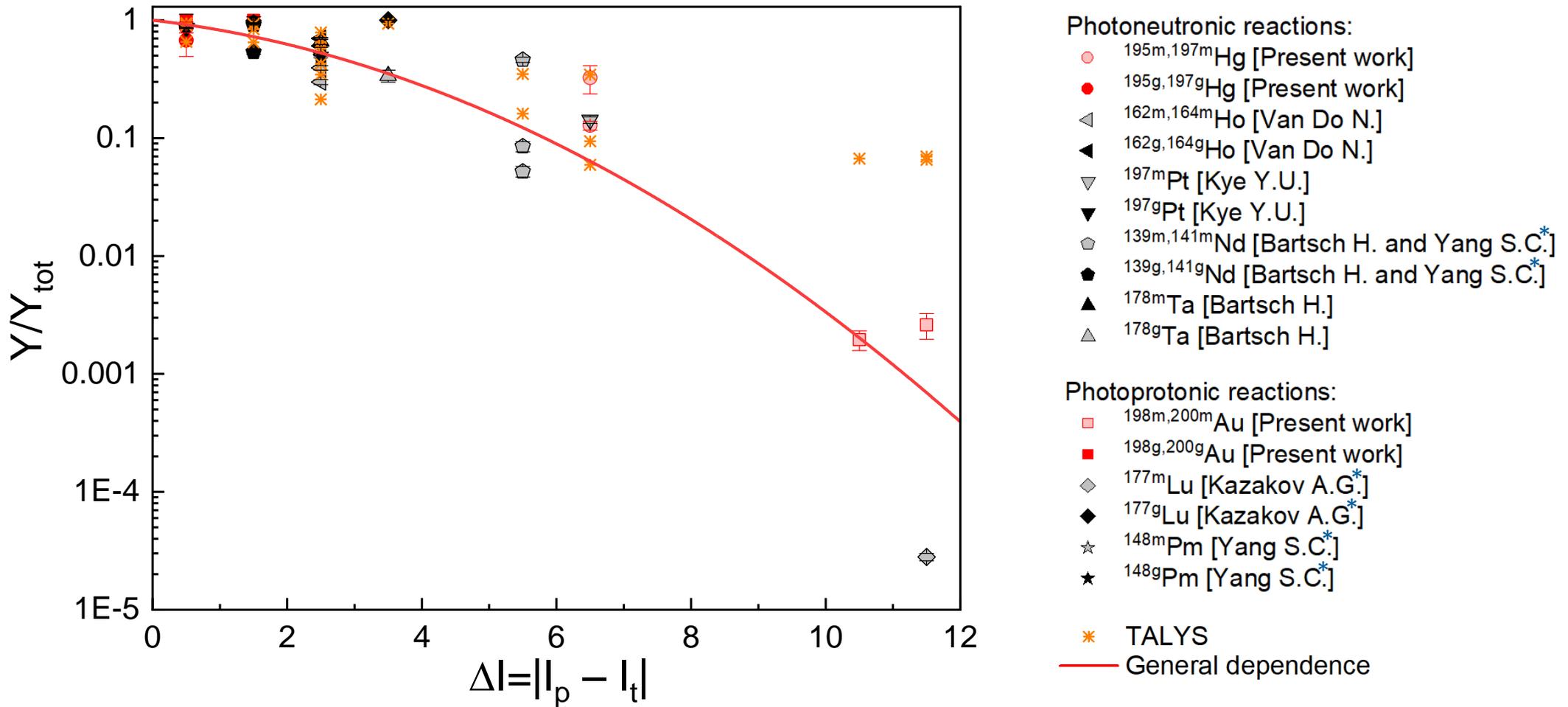
<i>Isotope</i>	<i>The main reaction of isotope production</i>	J_m^P	J_g^P	IR_{exp}	IR_{TALYS} (actual bremsstrahlung spectrum)
^{195}Hg	$^{196}\text{Hg}(0^+)(\gamma, 1n)^{195m,g}\text{Hg}$	$13/2^+$	$1/2^-$	0.481 ± 0.016	0.345
^{197}Hg	$^{198}\text{Hg}(0^+)(\gamma, 1n)^{197m,g}\text{Hg}$	$13/2^+$	$1/2^-$	0.147 ± 0.009	0.094
^{198}Au	$^{199}\text{Hg}(1/2^-)(\gamma, 1p)^{198m,g}\text{Au}$ $^{200}\text{Hg}(0^+)(\gamma, 1n1p)^{198m,g}\text{Au}$	12^-	2^-	0.0026 ± 0.0006	0.0576
^{200}Au	$^{201}\text{Hg}(3/2^-)(\gamma, 1p)^{200m,g}\text{Au}$ $^{202}\text{Hg}(0^+)(\gamma, 1n1p)^{200m,g}\text{Au}$	12^-	1^-	0.0019 ± 0.0003	0.0671

The **experimental** isomeric ratios for ^{195}Hg , ^{197}Hg are compared with **theoretical** calculations according to the **TALYS** program, as well as with works found in the literature



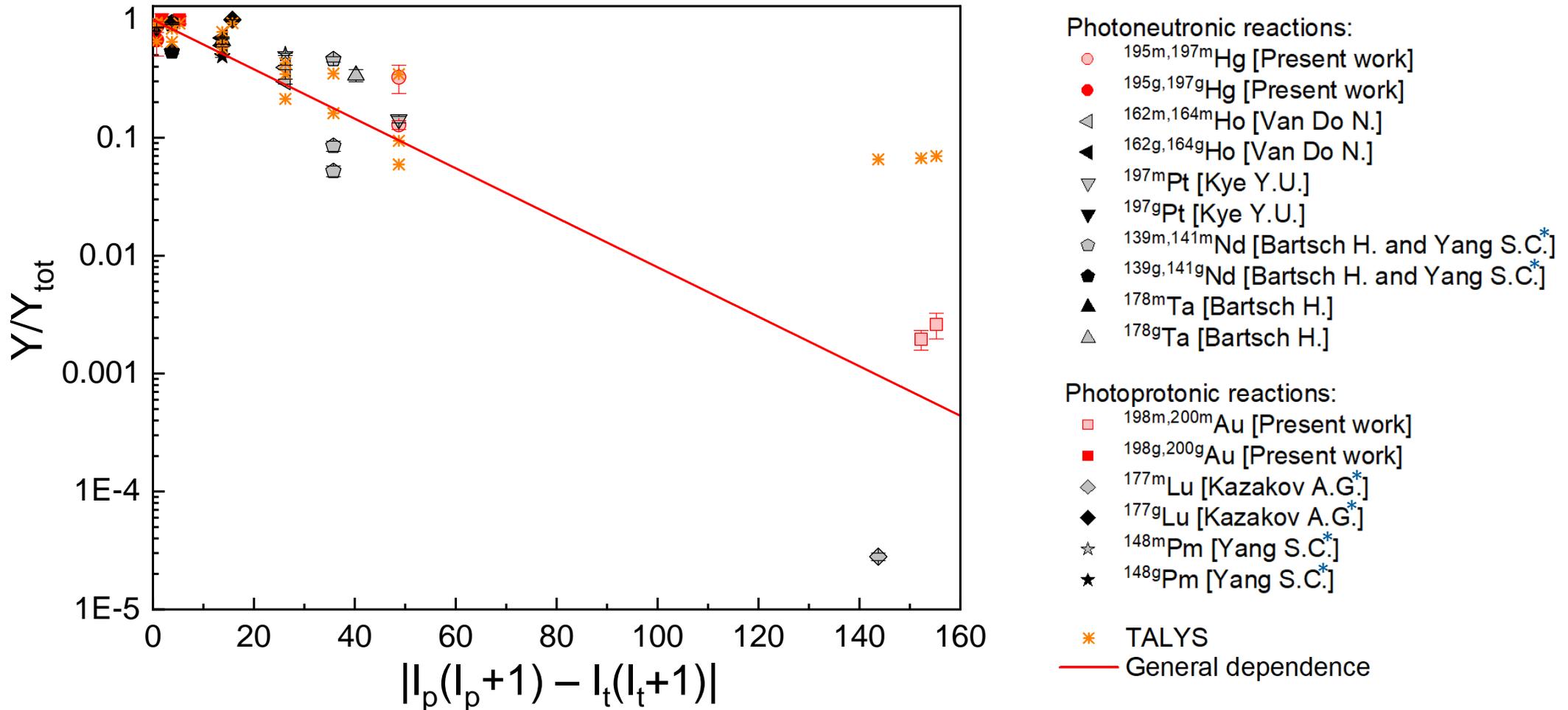
The dependence of the isomeric ratios of IR (^{195}Hg) (left graph) and IR (^{197}Hg) (right) on the irradiation energy

There is a clear **decrease in partial yields as the absolute spin difference increases**. Based on **experimental data**, **quadratic-exponential approximation curve** $Y/Y_{tot} = \exp(-A\Delta I^2 - B\Delta I)$ was fitted



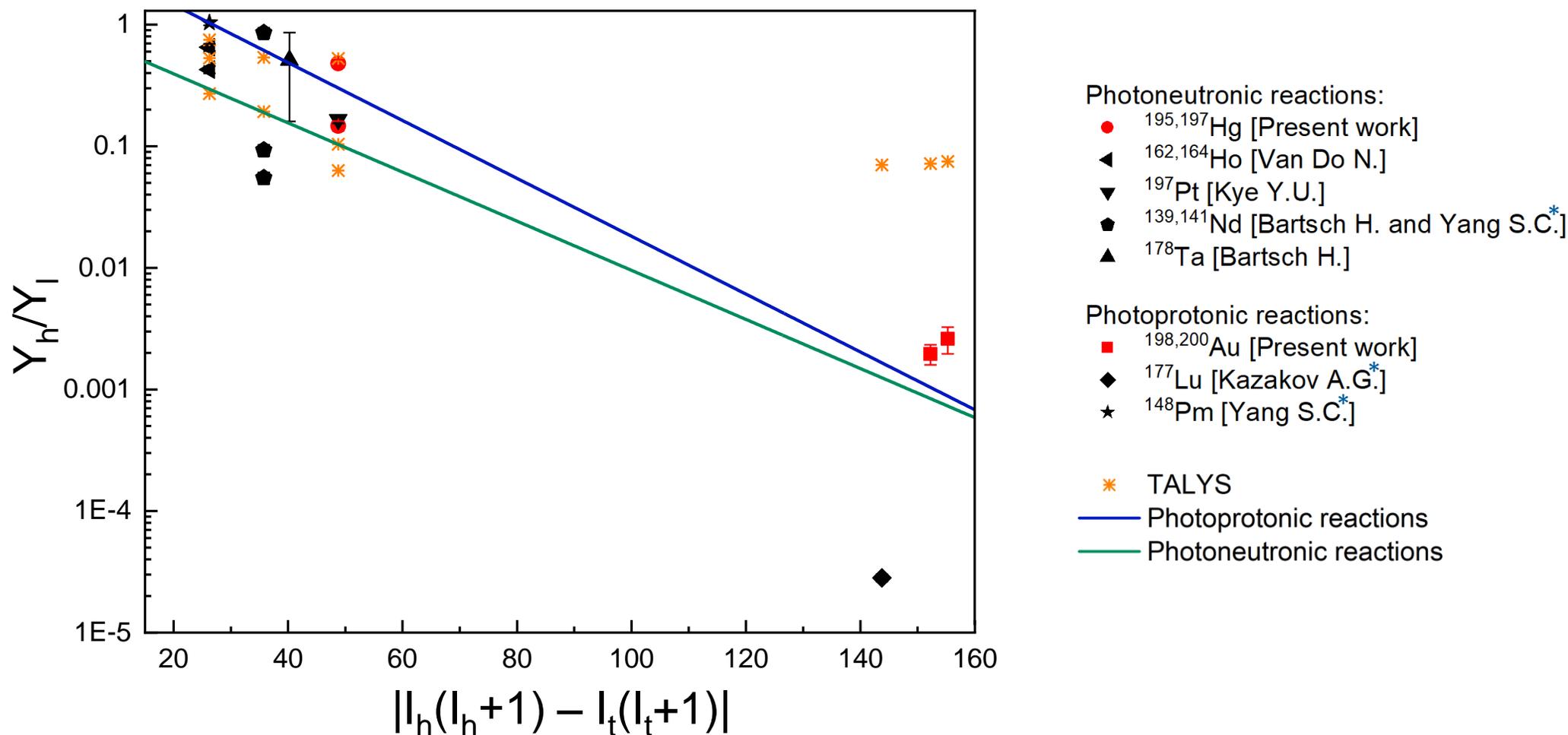
The dependence of partial yields of high-spin (lightly shaded points) and low-spin (fully shaded points) states **at an electron energy of 55 MeV** on the absolute value of the spin difference between the product and target $\Delta I = |I_p - I_t|$

Based on **experimental** data, **linear-exponential approximation curve** $Y/Y_{tot} = \exp(-A\Delta I)$ was fitted. The parameter $|I_p(I_p + 1) - I_t(I_t + 1)|$ shows **better correlation** with the population probabilities of metastable and ground states



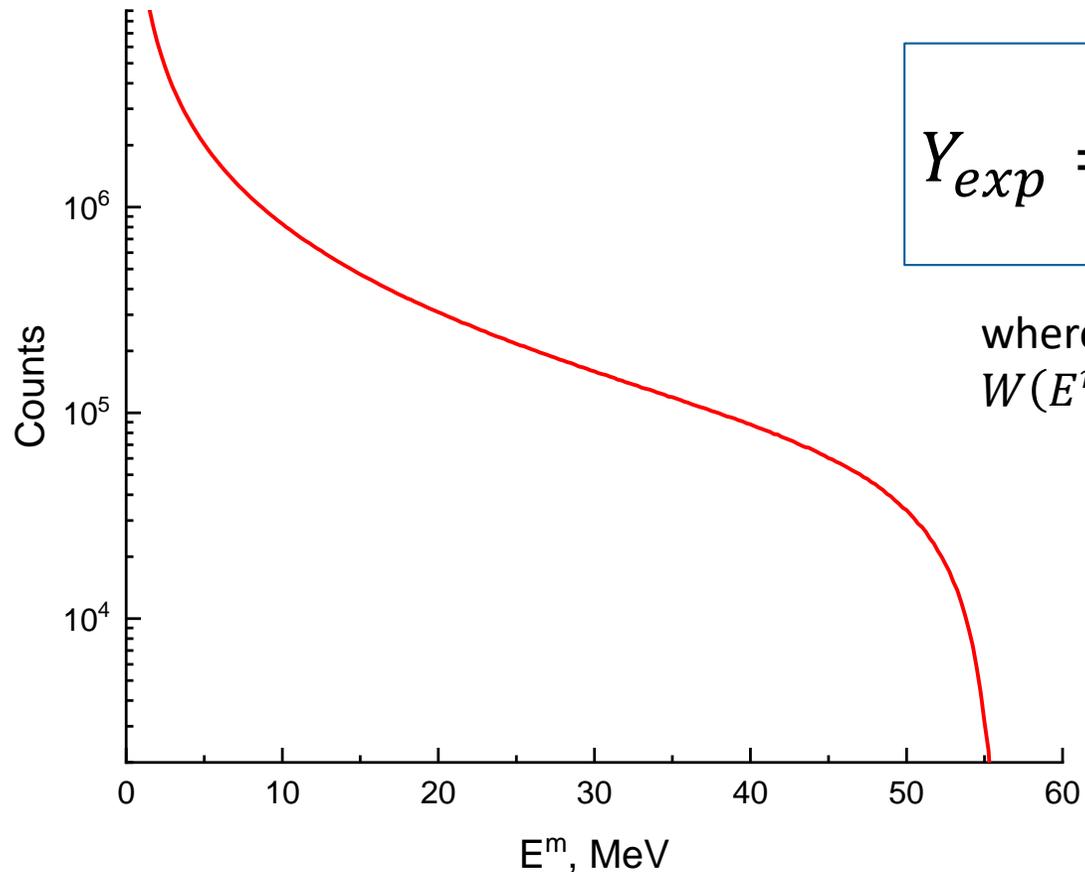
The dependence of partial yields of high-spin (lightly shaded points) and low-spin (fully shaded points) states **at an electron energy of 55 MeV** on the absolute value $\Delta I = |I_p(I_p + 1) - I_t(I_t + 1)|$

Under identical conditions, **the isomeric ratio $IR = Y_h/Y_l$ is consistently higher for photoproton reactions than for photoneutron reactions due to the Coulomb barrier effect**



The dependence of isomeric ratios **at an electron energy of 55 MeV** on the absolute value of the spin difference between the high-spin state and target $\Delta I = |I_h(I_h + 1) - I_t(I_t + 1)|$

A **final** conclusion on the existence of dependence common to all heavy nuclei **can be made only after additional studies of single-channel reactions either on enriched targets or at lower irradiation energies**



*Bremsstrahlung spectrum with an upper endpoint
energy of $E^m = 55$ MeV*

$$Y_{exp} = k_{norm} M \int_{E_{th}}^{E^m} W(E^m, E) \sigma(E) dE,$$

where M - the surface concentration of the target nuclei
 $W(E^m, E)$ - bremsstrahlung γ -ray spectrum

\Rightarrow **To reconstruct the reaction cross section $\sigma(E)$** , one must solve **a complex unfolding inverse problem** (employing the Penfold-Leiss approaches, Cook's least-structure method, Tikhonov regularization, among others).

The use of **monochromatic γ -ray beams** overcomes fundamental limitations of conventional sources and **enables direct measurements of nuclear reaction cross-sections**. The **NCPHM CRS** (Compton radiation source) represents such a facility

$$R = I_{\gamma} \cdot M \cdot d \cdot \sigma(E_{\gamma}),$$

where $I_{\gamma} \sim 5 \cdot 10^8 \text{ s}^{-1}$ – the intensity of gamma quanta, $M = \frac{\rho}{M_r} \cdot N_A$ – the surface concentration of the target nuclei, M_r – the molecular weight of the target, $d = 1 \text{ cm}$ – the target thickness, $\sigma(E_{\gamma})$ – the reaction cross-section at E_{γ} (calculated with TALYS)

Parameters		Reaction			
		$^{165}\text{Ho}(\gamma, 3n)^{162m}\text{Ho}$	$^{165}\text{Ho}(\gamma, 3n)^{162g}\text{Ho}$	$^{165}\text{Ho}(\gamma, 1n)^{164m}\text{Ho}$	$^{165}\text{Ho}(\gamma, 1n)^{164g}\text{Ho}$
$E_{th}, \text{ MeV}$		23.18	23.07	8.13	7.99
$\sigma(E_{\gamma}), \text{ mbar}$ (TALYS)	$E_{\gamma} = 10 \text{ MeV}$	0	0	$\cong 13.462$	$\cong 51.327$
	$E_{\gamma} = 20 \text{ MeV}$	0	0	$\cong 1.769$	$\cong 5.861$
	$E_{\gamma} = 30 \text{ MeV}$	$\cong 8.107$	$\cong 11.512$	$\cong 0.276$	$\cong 0.920$
$R, \text{ s}^{-1}$	$E_{\gamma} = 10 \text{ MeV}$	0	0	~ 216100	~ 824100
	$E_{\gamma} = 20 \text{ MeV}$	0	0	~ 28400	~ 94100
	$E_{\gamma} = 30 \text{ MeV}$	~ 130200	~ 184800	~ 4400	~ 14800

The total number of detected events from the NCPm CRS was evaluated, demonstrating the feasibility of the proposed measurement methodology

Parameters		Reaction			
		$^{165}\text{Ho}(\gamma, 3n)^{162m}\text{Ho}$	$^{165}\text{Ho}(\gamma, 3n)^{162g}\text{Ho}$	$^{165}\text{Ho}(\gamma, 1n)^{164m}\text{Ho}$	$^{165}\text{Ho}(\gamma, 1n)^{164g}\text{Ho}$
$E_\gamma, \text{keV} (I_\gamma, \%)$		282.86 (11.3%)	1319.3 (3.8%)	56.64 (6.5%)	91.40 (2.2%)
Detector efficiency $k (5 \text{ cm})$		0.012	0.006	0.0007	0.0014
$T_{1/2}$		67.1 min	15.0 min	36.6 min	28.8 min
N_{det} ($t = T_{1/2}$)	$E_\gamma = 10 \text{ MeV}$	0	0	~ 5	~ 13
	$E_\gamma = 20 \text{ MeV}$	0	0	~ 0.6	~ 1.4
	$E_\gamma = 30 \text{ MeV}$	~ 88	~ 21	~ 0.10	~ 0.23
Background N_f		~ 16	~ 0.4	~ 15	~ 14

Threshold and spin factors strongly influence the yields of the photonuclear reactions

1. **Experimental isomeric yield ratios** for $^{195,197}\text{Hg}$ and $^{198,200}\text{Au}$ at $E^m = 55 \text{ MeV}$ were calculated. Comparing the obtained IR with **the TALYS program** calculations indicates **acceptable agreement** for **photoneutron** reactions and **poor agreement** for **photoproton** reactions.
2. Possible **reasons for the discrepancies** between experimental and theoretical values include **theoretical parameters** and the use of the **bremstrahlung spectrum** calculated **based on the Seltzer-Berger tables instead of the actual** bremsstrahlung spectrum in thick targets.
3. The **theoretical** isomeric ratios calculated using the TALYS program reach **saturation**.
4. For many **heavy-nuclei** targets, **a systematic dependence of the yield on the spin factor has been observed**.
5. **The isomeric ratio** for **photoproton** reactions **is** systematically **higher** than for **photoneutron** reactions due to the presence of the **Coulomb barrier**.
6. The conducted evaluations demonstrate a fundamentally new capability to perform activation experiments and study isomeric states using the developed **the NCPHM CRS**.

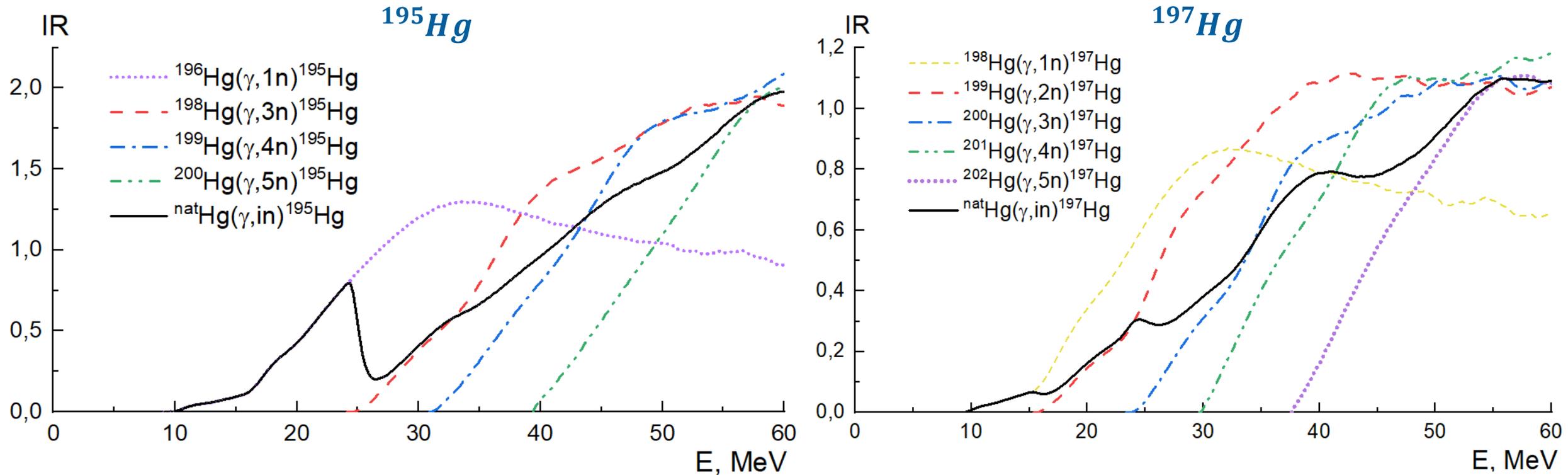
Isomeric ratios IR were obtained for **heavy-nuclei** isotopes produced in **photonuclear** reactions at the maximum electron accelerator energy of $E_m = 55 \text{ MeV}$

<i>Main reaction</i>	J_m^P	J_g^P	$IR_{exp} = Y_m/Y_g$	<i>Reference</i>
$^{196}\text{Hg}(0^+) (\gamma, 1n)^{195m,g}\text{Hg}$	$13/2^+$	$1/2^-$	0.481 ± 0.016	Present work*
$^{198}\text{Hg}(0^+) (\gamma, 1n)^{197m,g}\text{Hg}$	$13/2^+$	$1/2^-$	0.147 ± 0.009	Present work*
$^{199}\text{Hg}(3/2^-) (\gamma, 1p)^{198m,g}\text{Au}$ $^{200}\text{Hg}(0^+) (\gamma, 1n1p)^{198m,g}\text{Au}$	12^-	2^-	0.0026 ± 0.0006	Present work*
$^{201}\text{Hg}(3/2^-)(\gamma, 1p)^{200m,g}\text{Au}$ $^{202}\text{Hg}(0^+)(\gamma, 1n1p)^{200m,g}\text{Au}$	12^-	1^-	0.0019 ± 0.0003	Present work*
$^{178}\text{Hf}(0^+)(\gamma, 1p)^{177m,g}\text{Lu}$	$23/2^-$	$7/2^+$	$(28.2 \pm 2.0) \cdot 10^{-6}$	[A.G. Kazakov]*
$^{165}\text{Ho}(7/2^-)(\gamma, 1n)^{164m,g}\text{Ho}$	6^-	1^+	0.427 ± 0.029	[N. Van Do]
$^{165}\text{Ho}(7/2^-)(\gamma, 3n)^{162m,g}\text{Ho}$	6^-	1^+	0.652 ± 0.045	[N. Van Do]
$^{198}\text{Pt}(0^+)(\gamma, 1n)^{197m,g}\text{Pt}$	$13/2^+$	$1/2^-$	0.166 ± 0.012	[Y.U. Kye]
$^{149}\text{Sm}(7/2^-)(\gamma, 1p)^{148m,g}\text{Pm}$ $^{150}\text{Sm}(0^+)(\gamma, 1n1p)^{148m,g}\text{Pm}$	$5^-, 6^-$	1^-	1.035 ± 0.048	[S.C. Yang]*
$^{142}\text{Nd}(0^+)(\gamma, 1n)^{141m,g}\text{Nd}$	$11/2^-$	$3/2^+$	0.055 ± 0.006	[H. Bartsch]
$^{181}\text{Ta}(7/2^+)(\gamma, 3n)^{178m,g}\text{Ta}$	1^+	7^-	1.96 ± 0.35	[H. Bartsch]
$^{142}\text{Nd}(0^+)(\gamma, 1n)^{141m,g}\text{Nd}$ $^{143}\text{Nd}(7/2^-)(\gamma, 2n)^{141m,g}\text{Nd}$	$11/2^-$	$3/2^+$	0.093 ± 0.010	[S.C. Yang]*
$^{142}\text{Nd}(0^+)(\gamma, 3n)^{139m,g}\text{Nd}$ $^{143}\text{Nd}(7/2^-)(\gamma, 4n)^{139m,g}\text{Nd}$	$11/2^-$	$3/2^+$	0.859 ± 0.081	[S.C. Yang]*

An interesting feature: **the isomeric ratio for ^{195}Hg decreases sharply** in the region of electron energies around 24 MeV, followed by an increase

$$IR_{\text{nat}} = \frac{\sum_i \eta_i \sigma_{h_i}}{\sum_i \eta_i \sigma_{l_i}},$$

where σ_h , σ_l – cross sections of high-spin and low-spin state formation, η_i – abundance of the isotope in the natural mixture



The dependence of the isomeric-to-ground state formation cross section ratio for ^{195}Hg (left graph) and ^{197}Hg (right graph) on the γ -quanta energy, calculated using the TALYS program

Radioisotopes ^{198}Au и ^{199}Au are a promising **theranostic pair**

^{199}Au :

The average energy of the emitted β -particle is 84 keV ($\lambda \approx 100$ mm), which gives this radionuclide the ability to deliver energy, for example, to micrometastases and tumor cells near the surface of organs. Thus, it is well suited for radioimmunotherapy.

^{198}Au :

It is used in radiotherapy of various types of cancer. In recent years, it has been used as gold nanoparticles for the development of radiopharmaceuticals.

Nanostructures labeled ^{198}Au or ^{199}Au are also being investigated as a way to visualize oncological diseases in vivo.

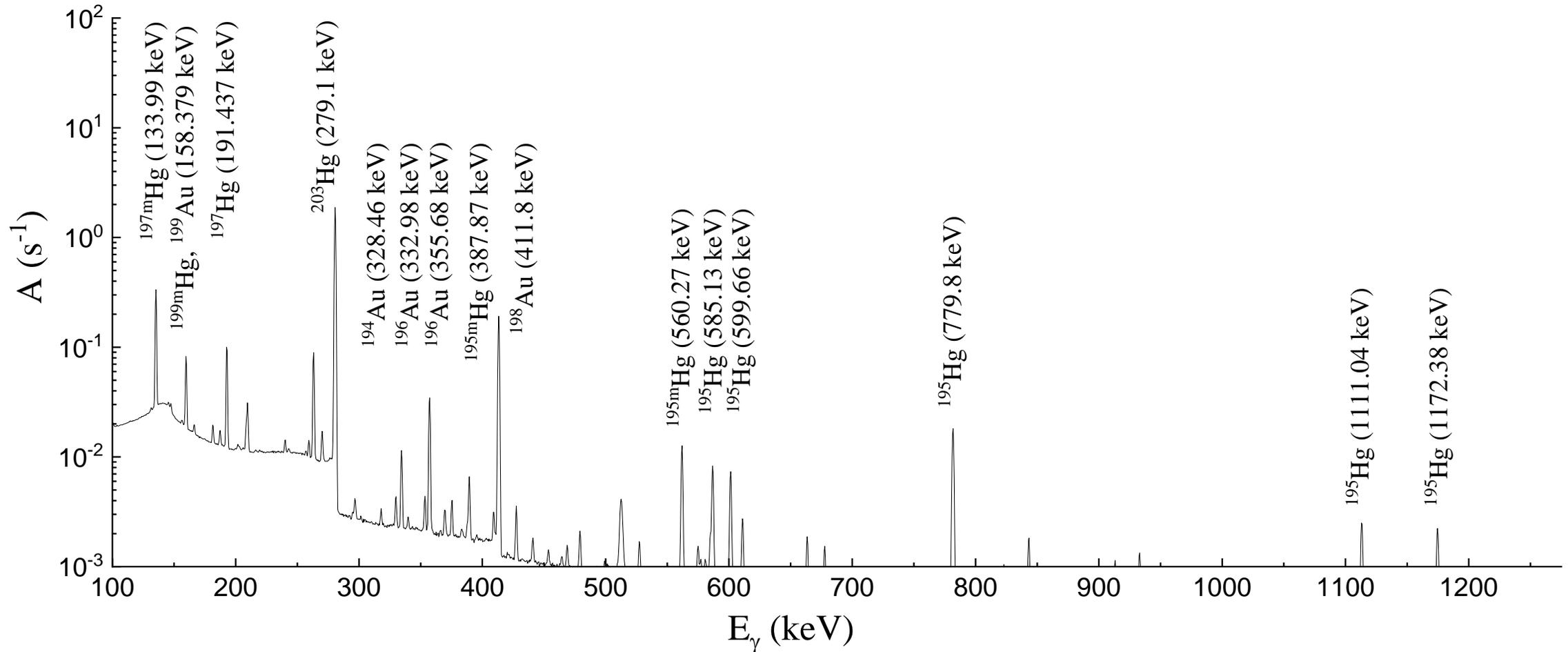
Various decay channels were identified by **the energy of the peaks E_γ in the spectrum** and **the half-lives $T_{1/2}$** of the isotopes formed

<i>Isotope</i>	<i>$T_{1/2}$</i>	<i>Isotope production reaction</i>	<i>$E_\gamma, keV (I_\gamma, \%)$</i>
^{195g}Hg	10.53 h	$^{nat}\text{Hg}(\gamma, in) = 0.00155 \cdot ^{196}\text{Hg}(\gamma, 1n) + 0.0997 \cdot ^{198}\text{Hg}(\gamma, 3n) + 0.1687 \cdot ^{199}\text{Hg}(\gamma, 4n) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 5n)$	180.11 (1.9), 207.1 (1.57), 261.75 (1.5), 585.13 (1.99), 599.66 (1.78), 779.8 (7.0), 1111.04 (1.44), 1172.38 (1.24)
^{195m}Hg	41.6 h	$^{nat}\text{Hg}(\gamma, in) = 0.00155 \cdot ^{196}\text{Hg}(\gamma, 1n) + 0.0997 \cdot ^{198}\text{Hg}(\gamma, 3n) + 0.1687 \cdot ^{199}\text{Hg}(\gamma, 4n) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 5n)$	261.75 (30.9), 387.87 (2.15), 560.27 (7.0)
^{197g}Hg	64.14 h	$^{nat}\text{Hg}(\gamma, in) = 0.0997 \cdot ^{198}\text{Hg}(\gamma, 1n) + 0.1687 \cdot ^{199}\text{Hg}(\gamma, 2n) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 3n) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 4n) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 5n)$	77.351 (18.7), 191.437 (0.632), 268.78 (0.04)
^{197m}Hg	23.8 h	$^{nat}\text{Hg}(\gamma, in) = 0.0997 \cdot ^{198}\text{Hg}(\gamma, 1n) + 0.1687 \cdot ^{199}\text{Hg}(\gamma, 2n) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 3n) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 4n) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 5n)$	133.99 (33.0), 164.97 (0.26), 279.197 (6.0)
^{199m}Hg	42.6 m	$^{nat}\text{Hg}(\gamma, in) = 0.2310 \cdot ^{200}\text{Hg}(\gamma, 1n) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 2n) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 3n)$	158.3795 (52.0), 374.1 (13.8)
^{203}Hg	46.612 d	$0,0687 \cdot ^{204}\text{Hg}(\gamma, 1n)$	279.1967 (81.0)

Various decay channels were identified by **the energy of the peaks E_γ in the spectrum** and **the half-lives $T_{1/2}$** of the isotopes formed

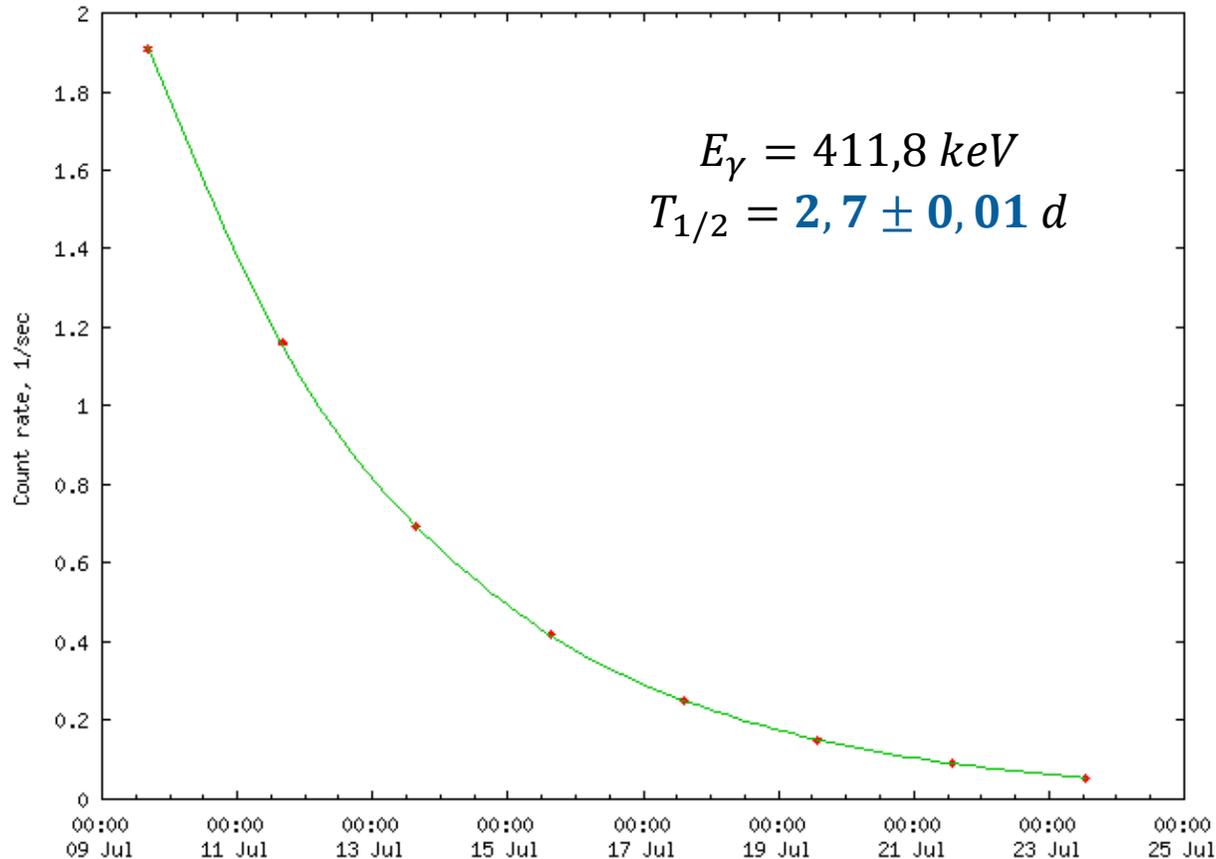
<i>Isotope</i>	<i>$T_{1/2}$</i>	<i>Isotope production reaction</i>	<i>$E_\gamma, keV (I_\gamma, \%)$</i>
^{194}Au	38.02 h	$^{nat}\text{Hg}(\gamma, in1p) = 0,00155 \cdot ^{196}\text{Hg}(\gamma, 1n1p) + 0,0997 \cdot ^{198}\text{Hg}(\gamma, 3n1p) + 0,1687 \cdot ^{199}\text{Hg}(\gamma, 4n1p)$	328.455 (61.0)
^{196}Au	6.183 d	$^{nat}\text{Hg}(\gamma, in1p) = 0.0997 \cdot ^{198}\text{Hg}(\gamma, 1n1p) + 0.1687 \cdot ^{199}\text{Hg}(\gamma, 2n1p) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 3n1p)$	332.983 (22.9), 355.684 (87.0)
^{198g}Au	2.695 d	$^{nat}\text{Hg}(\gamma, in1p) = 0.1687 \cdot ^{199}\text{Hg}(\gamma, 1p) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 1n1p) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 2n1p)$	411.802 (96.0)
^{198m}Au	2.27 d	$^{nat}\text{Hg}(\gamma, in1p) = 0.1687 \cdot ^{199}\text{Hg}(\gamma, 1p) + 0.2310 \cdot ^{200}\text{Hg}(\gamma, 1n1p) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 2n1p)$	214.841 (77.0)
^{199}Au	3.139 d	$^{nat}\text{Hg}(\gamma, in1p) = 0.2310 \cdot ^{200}\text{Hg}(\gamma, 1p) + 0.1318 \cdot ^{201}\text{Hg}(\gamma, 1n1p) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 2n1p)$	158.3795 (40.0), 208.206 (8.73)
^{200g}Au	48.4 m	$^{nat}\text{Hg}(\gamma, in1p) = 0.1318 \cdot ^{201}\text{Hg}(\gamma, 1p) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 1n1p)$	367.943 (19.0), 1225.479 (10.7), 1262.950 (3.12)
^{200m}Au	18.7 h	$^{nat}\text{Hg}(\gamma, in1p) = 0.1318 \cdot ^{201}\text{Hg}(\gamma, 1p) + 0.2986 \cdot ^{202}\text{Hg}(\gamma, 1n1p)$	255.87 (71.0), 367.943 (73.0)

The yields of photonuclear reactions were determined from **the peaks of the γ -lines in the spectra of the residual activity** of the irradiated mercury sample



The spectrum of residual activity of the irradiated Hg sample (data set was carried out for 16 days)

The sources of the peaks were identified using **The Lund/LBNL Nuclear Data database** and **the determination of the half-life $T_{1/2}$**



$T_{1/2}$ were identified using an automatic acquisition system.

Gammas from ^{198}Au (2.69517 d 21)

E_γ (keV)	I_γ (%)	Decay mode
411.80205	17 96	β^-
675.8836 7	0.804 3	β^-
1087.684 3	0.159 2	β^-

Thus, the peak of 411.8 keV corresponds to the **pure** decay channel ^{198}Au

The half-life determined by approximating a decrease in the peak area corresponding to an energy of 411.8 keV

The results of **theoretical** calculations of the cross sections of photonuclear reactions on Hg isotopes using **the combined photonucleon reaction model** (CMPR) were analyzed

<i>Isotope</i>	<i>Isotope production reaction</i>	$E_{th}, \text{MэВ}$	$Y(55 \text{ MэВ})_{exp}, e^{-1}$	Y_{CMPR}, e^{-1}	Y_{TALYS}, e^{-1}
$^{195g+m}\text{Hg}$	$^{nat}\text{Hg}(\gamma, in) =$ $0.0015 \cdot ^{196}\text{Hg}(\gamma, 1n)$ $+ 0.0997 \cdot ^{198}\text{Hg}(\gamma, 3n)$ $+ 0.1687 \cdot ^{199}\text{Hg}(\gamma, 4n)$	8.90 24.17 30.83	$(3.8 \pm 1.0) \cdot 10^{-6}$	$3.992 \cdot 10^{-6}$	$5.516 \cdot 10^{-6}$
$^{197g+m}\text{Hg}$	$^{nat}\text{Hg}(\gamma, in) =$ $0.0997 \cdot ^{198}\text{Hg}(\gamma, 1n)$ $+ 0.1687 \cdot ^{199}\text{Hg}(\gamma, 2n)$ $+ 0.2310 \cdot ^{200}\text{Hg}(\gamma, 3n)$	8.45 15.15 23.18	$(1.00 \pm 0.07) \cdot 10^{-4}$	$0.982 \cdot 10^{-4}$	$1.025 \cdot 10^{-4}$
^{199m}Hg	$^{nat}\text{Hg}(\gamma, in) =$ $0.2310 \cdot ^{200}\text{Hg}(\gamma, 1n)$ $+ 0.1318 \cdot ^{201}\text{Hg}(\gamma, 2n)$ $+ 0.2986 \cdot ^{202}\text{Hg}(\gamma, 3n)$	8.56 14.79 22.54	$(1.56 \pm 0.09) \cdot 10^{-5}$	-*	$8.417 \cdot 10^{-6}$
^{203}Hg	$^{nat}\text{Hg}(\gamma, in) =$ $0.0687 \cdot ^{204}\text{Hg}(\gamma, 1n)$	7.49	$(4.97 \pm 0.25) \cdot 10^{-5}$	$4.587 \cdot 10^{-5}$	$4.081 \cdot 10^{-5}$

There is a **good agreement** of the experimental data within the error limits with the results of the calculation according to **the CMPR** in the case of **photoproton** reactions, according to **the TALYS program, discrepancies are observed**

<i>Isotope</i>	<i>Isotope production reaction</i>	$E_{th}, \text{MэВ}$	$Y(55 \text{ MэВ})_{exp}, e^{-1}$	Y_{CMPR}, e^{-1}	Y_{TALYS}, e^{-1}
^{194}Au	$^{nat}\text{Hg}(\gamma, in1p) =$ $0.0015 \cdot ^{196}\text{Hg}(\gamma, 1n1p)$ $+ 0.0997 \cdot ^{198}\text{Hg}(\gamma, 3n1p)$	14.93 30.24	$(1.74 \pm 0.11) \cdot 10^{-8}$	$8.530 \cdot 10^{-9}$	$9.781 \cdot 10^{-9}$
^{196}Au	$^{nat}\text{Hg}(\gamma, in1p) =$ $0.0997 \cdot ^{198}\text{Hg}(\gamma, 1n1p)$ $+ 0.1687 \cdot ^{199}\text{Hg}(\gamma, 2n1p)$	15.18 21.84	$(2.34 \pm 0.12) \cdot 10^{-7}$	$2.000 \cdot 10^{-7}$	$1.577 \cdot 10^{-7}$
$^{198g+m}\text{Au}$	$^{nat}\text{Hg}(\gamma, in1p) =$ $0.1687 \cdot ^{199}\text{Hg}(\gamma, 1p)$ $+ 0.2310 \cdot ^{200}\text{Hg}(\gamma, 1n1p)$	7.25 15.28	$(1.11 \pm 0.06) \cdot 10^{-6}$	$1.329 \cdot 10^{-6}$	$3.468 \cdot 10^{-7}$
^{199}Au	$^{nat}\text{Hg}(\gamma, in1p) =$ $0.2310 \cdot ^{200}\text{Hg}(\gamma, 1p)$ $+ 0.1318 \cdot ^{201}\text{Hg}(\gamma, 1n1p)$	7.70 13.93	$(1.34 \pm 0.07) \cdot 10^{-6}$	$1.384 \cdot 10^{-6}$	$3.382 \cdot 10^{-7}$
$^{200g+m}\text{Au}$	$^{nat}\text{Hg}(\gamma, in1p) =$ $0.1318 \cdot ^{201}\text{Hg}(\gamma, 1p)$ $+ 0.2986 \cdot ^{202}\text{Hg}(\gamma, 1n1p)$	7.68 15.44	$(9.2 \pm 0.8) \cdot 10^{-7}$	$8.864 \cdot 10^{-7}$	$2.757 \cdot 10^{-7}$