

LXXV International Conference NUCLEUS 2025
St. Petersburg, Russia July 1st – 6th, 2025

Nuclear Collective Vibrations Studied by Beyond Mean-Field Approach

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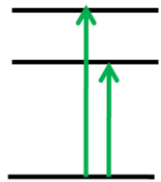


- **Nuclear collective vibrations**
- **Collective vibrations provide insight to**
 - ✓ **How were the heavy elements from iron to uranium made?**
 - ✓ **What is the equation of state (EOS) of nuclear matter?**
- **New possibilities to explore collective vibrations**
- **Summary and Perspective**

Nuclear Collective Vibrations

➤ The vibration excitation of nucleus involving many nucleons

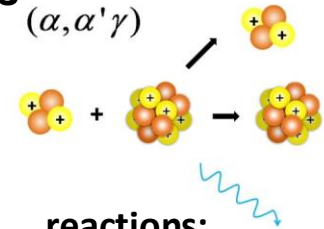
• Non-charge-exchange excitations



Z, N

- ✓ Giant monopole resonances (GMR)
- ✓ Giant dipole resonances (GDR)
- ✓ Giant quadrupole resonances (GQR)

...

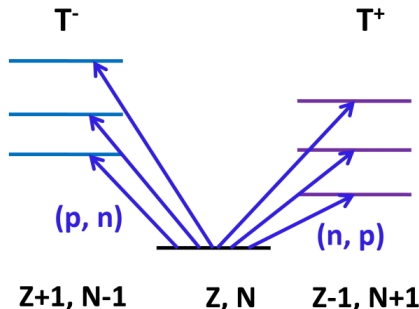


reactions:

$(\alpha, \alpha' \gamma)$,
 (γ, γ') , (γ, n) , (e, e')

Strong
EM

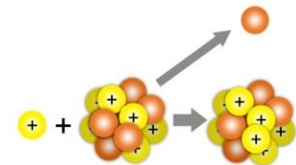
• Charge-exchange excitations



- ✓ Isobaric Analogue States (IAS)
- ✓ Gamow-Teller Resonances (GTR)
- ✓ Spin-Dipole (SD) excitations

...

(p,n) reaction



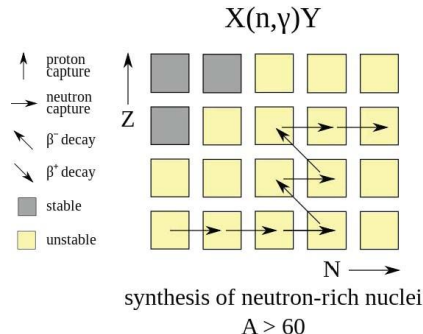
reactions:

$(p,n) T^-$, $(n,p) T^+$

Strong

Collective Vibrations Provide Insight to

- 🧐 ***How were the heavy elements from iron to uranium made?***
- 🧐 ***What is the equation of state (EOS) of nuclear matter?***
- 🧐 ***How do stars explode?***
- 🧐 ***What are the masses of neutrinos?***
- 🧐 ***What is the interaction between nucleons in nuclear medium that governs the properties of nuclei?***

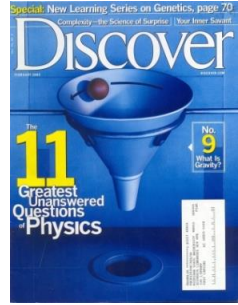


How Were the Heavy Elements Made?

How were the heavy elements from iron to uranium made?

- r-process

The 11 greatest unanswered questions of physics



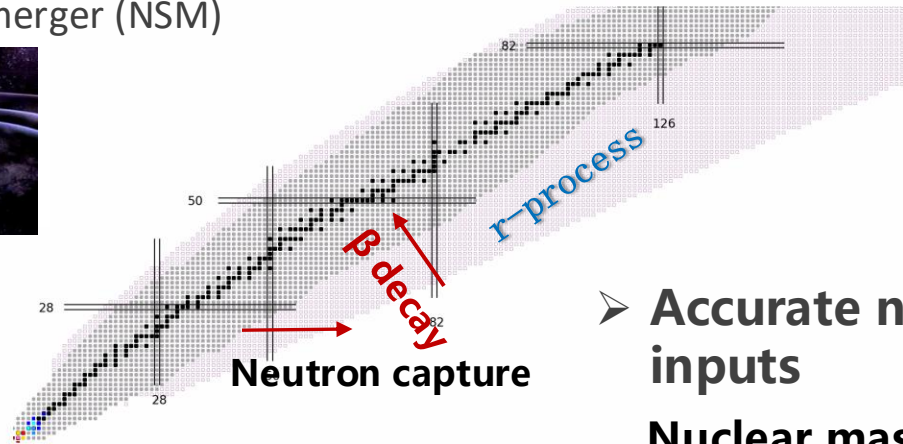
- Where does r-process happen?

Supernova Neutron star merger (NSM)



GW170817 NSM:
One of the main r-process
sites

Nature **551**,64; 67; 75; 80 (2017)
Science **358**, 1559 (2017)
ApJL **848**, L17; L19 (2017)



**R-process path: far from
stability, relies on theory!**

- **Accurate nuclear physics
inputs**

Nuclear mass,

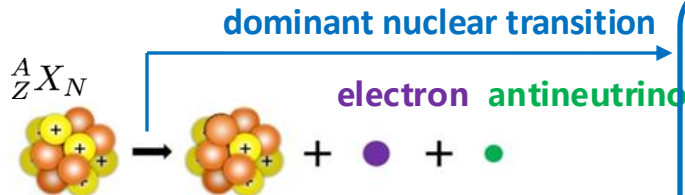
β decay half-lives,

Neutron-capture rates,

...

β -decay Half-life Is a Hard Problem

- β -decay



Gamow-Teller transition

$$T^- \quad \Delta S=1 \quad \Delta L=0 \quad \Delta T=1$$

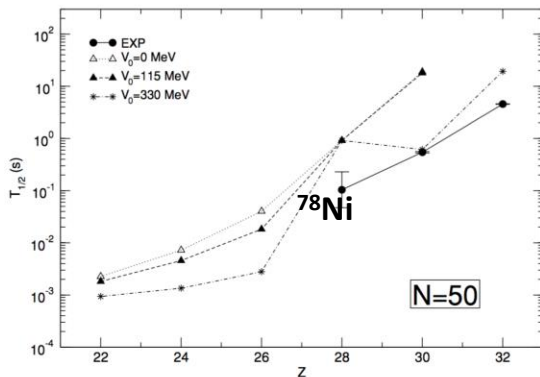
Charge-exchange mode

$$\hat{O}_{GT^-} = \sum_{i=1}^A \vec{\sigma}(i) \cdot \tau_-(i)$$

$Z+1, N-1 \quad Z, N$

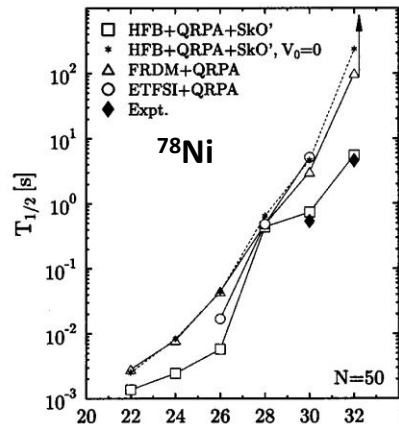
- RHB+QRPA

(quasiparticle random phase approximation)



Niksic, et al., PRC 71, 014308 (2005)

- Skyrme HFB+QRPA



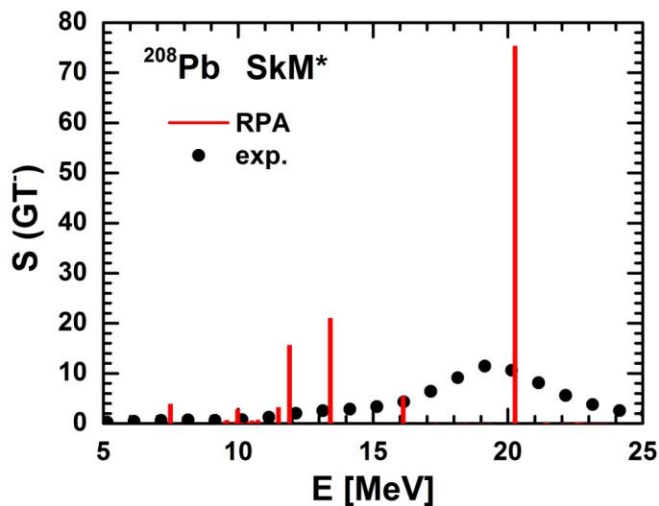
Engel, et al., PRC 60, 014302 (1999)

- Half-lives are **overestimated**

✓ Due to the nuclear structure part – Gamow-Teller transition

Limitations of (Q)RPA Description

- (Q)RPA cannot describe the spreading width



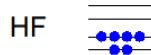
- Spreading Width (Damping Width)

energy and angular momentum of coherent vibrations

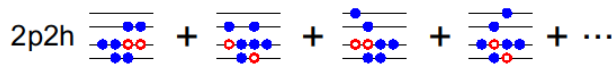
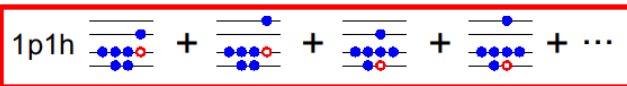
→ more complicated states of 2p-2h, 3p-3h, ... character

- Correlations beyond RPA

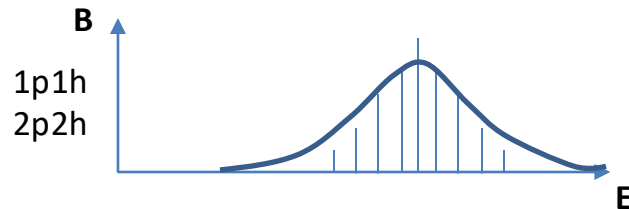
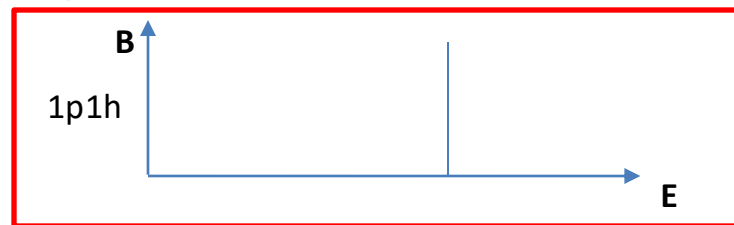
HF



RPA

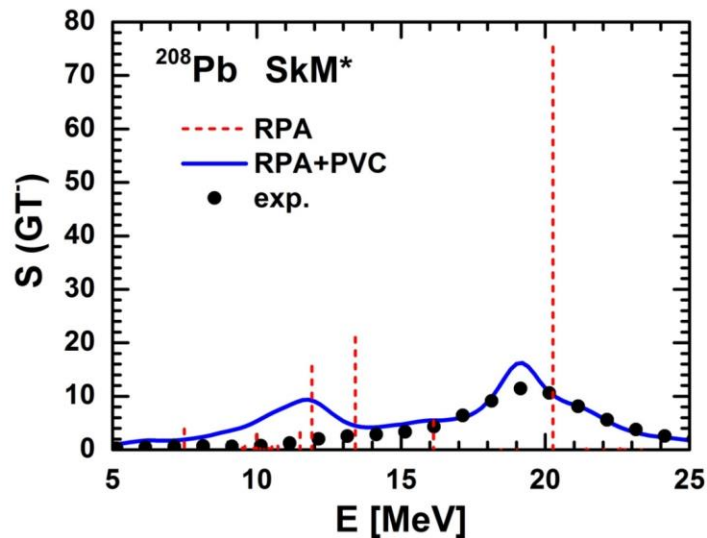


⋮



Limitations of (Q)RPA Description

- Particle Vibration Coupling (PVC) effect



- ✓ Develop a spreading width
- ✓ Reproduce resonance lineshape

Y. F. Niu, G. Colo, and E. Vigezzi, *PRC* 90, 054328 (2014)

Y. F. Niu, G. Colo, and E. Vigezzi, *PRC* 94, 064328 (2016)

- Correlations beyond RPA

HF

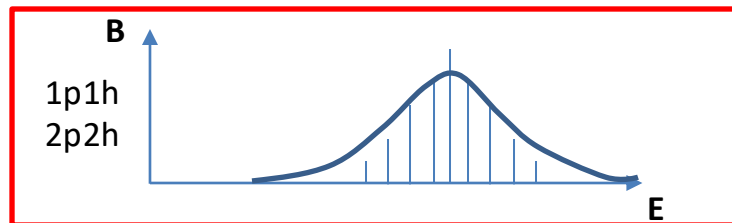
RPA

1p1h

~PVC

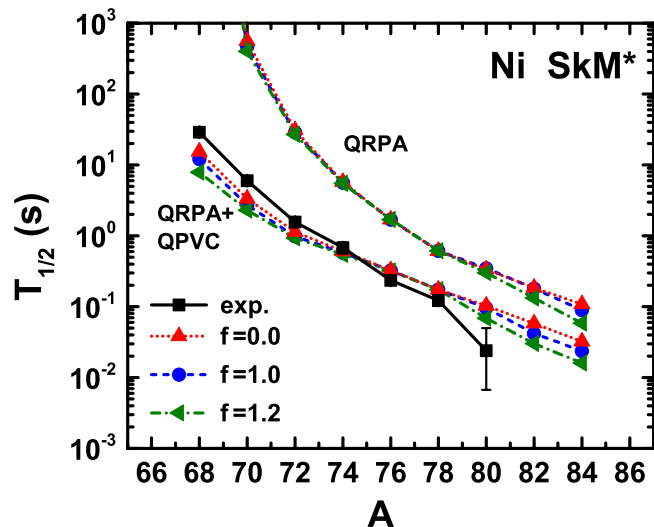
2p2h

⋮

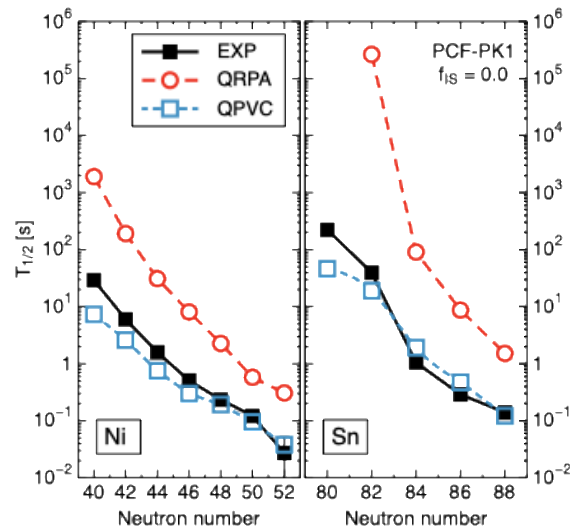


β -Decay Half-Lives in Ni and Sn isotopes

- Skyrme Quasiparticle PVC (QPVC)



- Relativistic Quasiparticle PVC (QPVC)

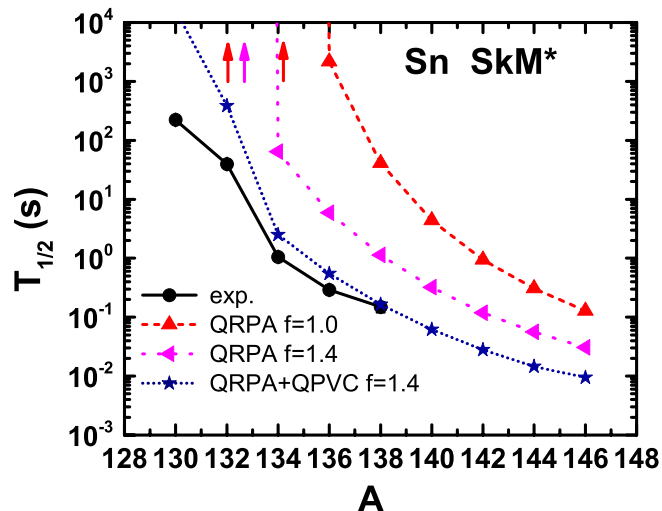


- Isoscalar Pairing:
 - similar at QRPA and QRPA+QPVC level
 - not effective for Ni isotopes (nuclei before N=50 closed shell)

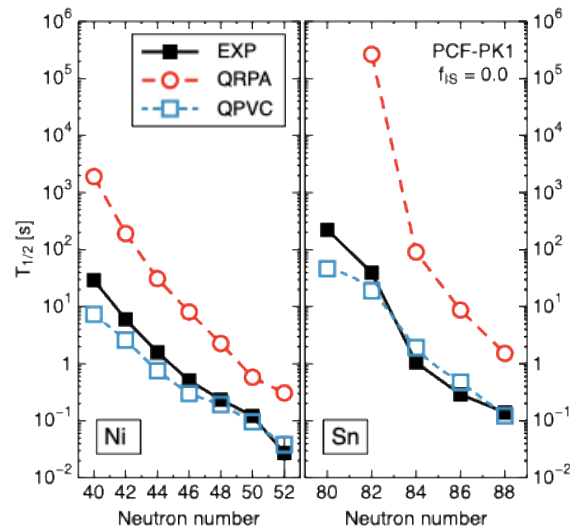
- QPVC:
 - reduce the half-lives

β -Decay Half-Lives in Ni and Sn isotopes

- Skyrme Quasiparticle PVC (QPVC)



- Relativistic Quasiparticle PVC (QPVC)



- Isoscalar Pairing:
 - similar at QRPA and QRPA+QPVC level
 - effective for Sn isotopes (nuclei above N=82 closed shell)

- QPVC:
 - reduce the half-lives

Collective Vibrations Provide Insight to

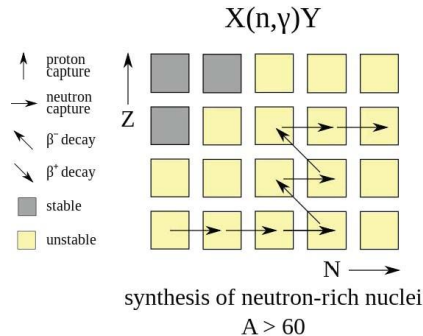
🧐 *How were the heavy elements from iron to uranium made?*

🧐 **What is the equation of state (EOS) of nuclear matter?**

🧐 *How do stars explode?*

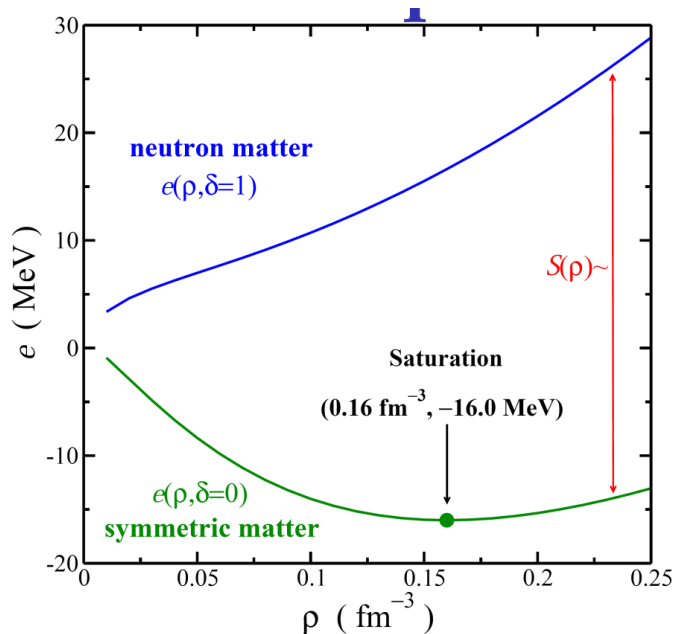
🧐 *What are the masses of neutrinos?*

🧐 *What is the interaction between nucleons in nuclear medium that governs the properties of nuclei?*



What Is the Equation of State of Nuclear Matter?

➤ Nuclear equation of state (EoS)



B. A. Li, L. W. Chen, and C. M. Ko,
Phys. Rep. 464, 113 (2008)

Nuclear matter
(NM)

$$\frac{E}{A}(\rho, \delta)$$

Symmetric
NM

$$\frac{E}{A}(\rho, \delta = 0)$$

symmetry
energy

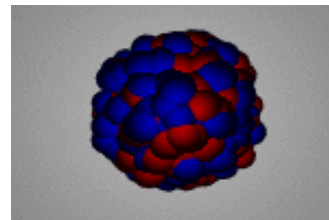
$$+ S(\rho)\delta^2$$

$$\delta \equiv (\rho_n - \rho_p)/\rho$$

$$\xi = \frac{\rho - \rho_0}{3\rho_0}$$

$$\frac{E}{A}(\rho, \delta = 0) = \epsilon_0 + \frac{1}{2}K\xi^2 + o(\xi^3)$$

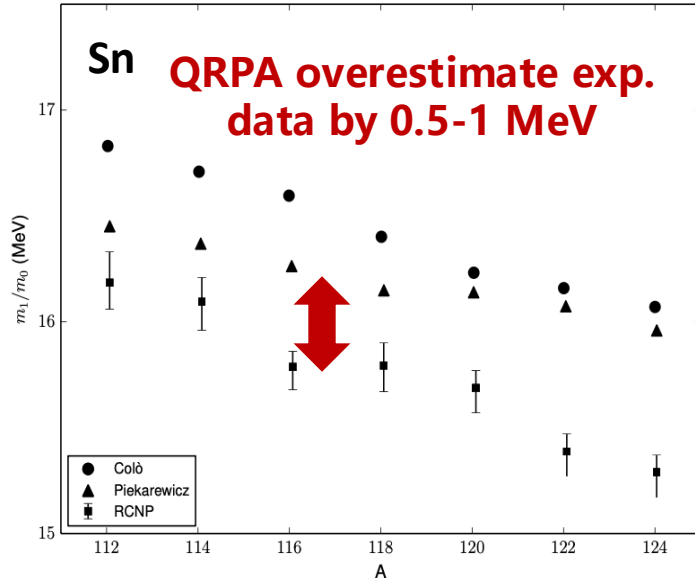
Nuclear incompressibility



**Giant Monopole
Resonance (GMR)**

**Non-charge-
exchange
mode**

Puzzle: Why are tins so soft?



Focus issue on open problems in
nuclear structure theory

J. Piekarewicz, **JPG** 37, 064038 (2010)

PRL 99, 162503 (2007)

PHYSICAL REVIEW LETTERS

week ending
19 OCTOBER 2007

Isotopic Dependence of the Giant Monopole Resonance in the Even- A $^{112-124}\text{Sn}$ Isotopes and the Asymmetry Term in Nuclear Incompressibility

T. Li,¹ U. Garg,¹ Y. Liu,¹ R. Marks,¹ B. K. Nayak,¹ P. V. Madhusudhana Rao,¹ M. Fujiwara,² H. Hashimoto,² K. Kawase,² K. Nakanishi,² S. Okumura,² M. Yosoi,² M. Itoh,³ M. Ichikawa,³ R. Matsuo,³ T. Terazono,³ M. Uchida,⁴ T. Kawabata,⁵ H. Akimune,⁶ Y. Iwao,⁷ T. Murakami,⁷ H. Sakaguchi,⁷ S. Terashima,⁷ Y. Yasuda,⁷ J. Zenihiro,⁷ and M. N. Harakeh⁸

The strength distributions of the giant monopole resonance (GMR) have been measured in the even- A Sn isotopes ($A = 112-124$) with inelastic scattering of 400-MeV α particles in the angular range $0^\circ - 8.5^\circ$. We find that the experimentally observed GMR energies of the Sn isotopes are lower than the values predicted by theoretical calculations that reproduce the GMR energies in ^{208}Pb and ^{90}Zr very well. From the GMR data, a value of $K_\tau = -550 \pm 100$ MeV is obtained for the asymmetry term in the nuclear incompressibility.

In Sn the ISGMR centroid energy is overestimated by about 1 MeV by the same models which reproduce the ISGMR energy well in ^{208}Pb .

**^{208}Pb GMR $\rightarrow$ K $\sim$ 230 MeV
v.s.
Sn GMR $\rightarrow$ K $\sim$ 205 MeV**

Garg and Colo, **PPNP** 101, 55 (2018)

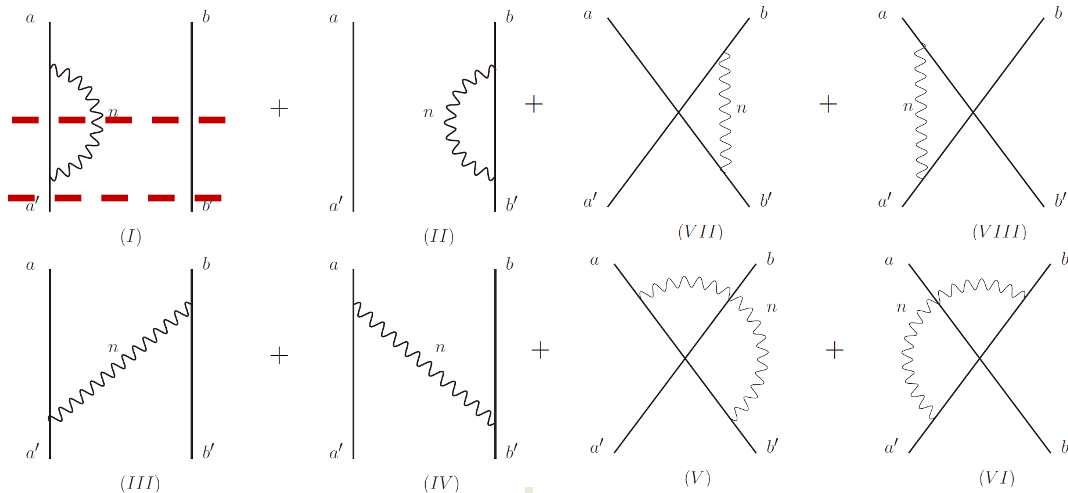
Our attempt to solve the puzzle

- Quasiparticle RPA + quasiparticle vibration coupling
(QRPA) + **(QPVC)** for non-charge-exchange modes

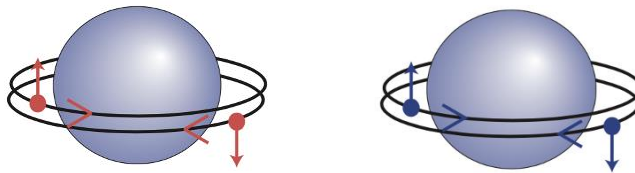
✓ QPVC effect

2 qp $\otimes$ phonons

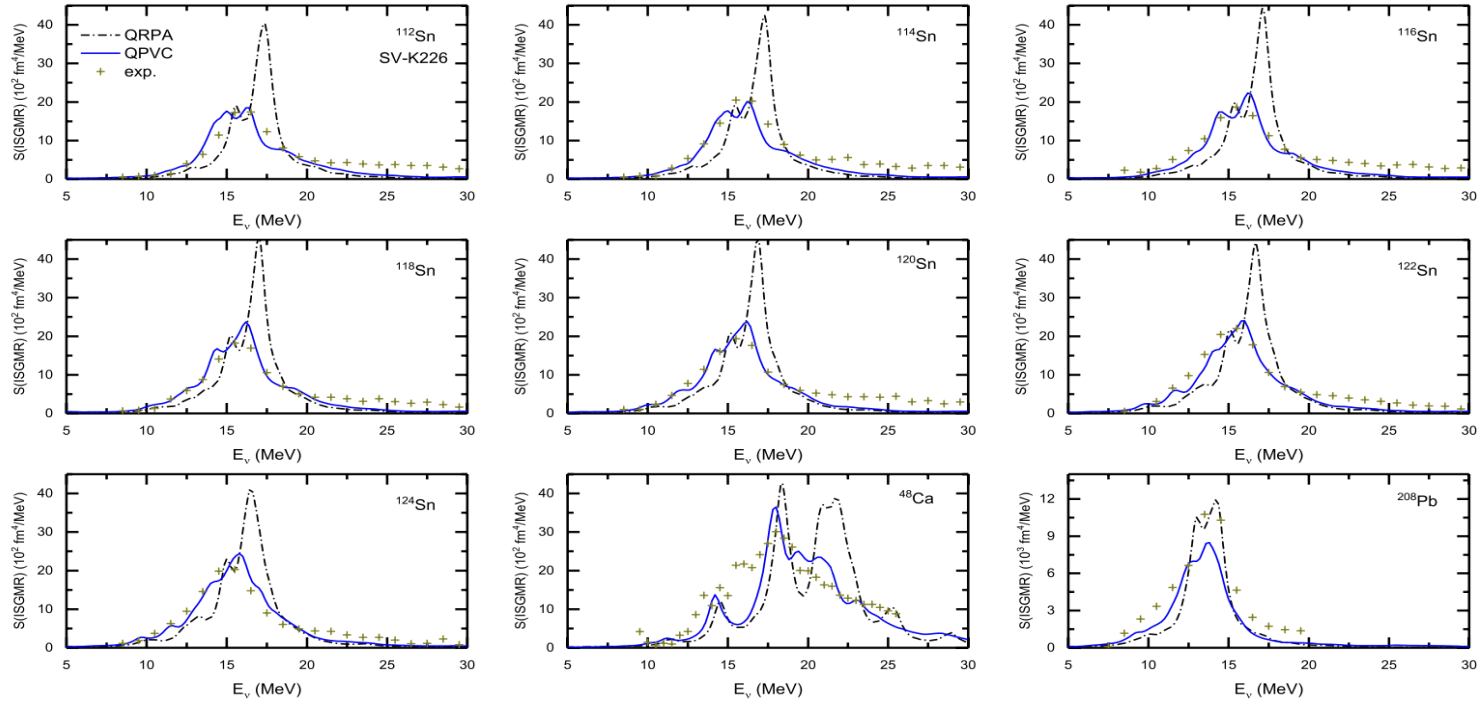
2 quasiparticles(qp)



✓ Pairing effect



Puzzle Solved: Unified description of GMR in Ca, Sn and Pb



SV-K226: $K_\infty = 226 \text{ MeV}$

Z. Z. Li, Y. F. Niu, and G. Colo, Phys. Rev. Lett. **131**,
082501, 2023

Outline

- Nuclear collective vibrations
- Collective vibrations provide insight to
 - ✓ How were the heavy elements from iron to uranium made?
 - ✓ What is the equation of state (EOS) of nuclear matter?
- **New possibilities to explore collective vibrations**
- Summary and Perspective

Photo-nuclear reaction

- **Advantages**

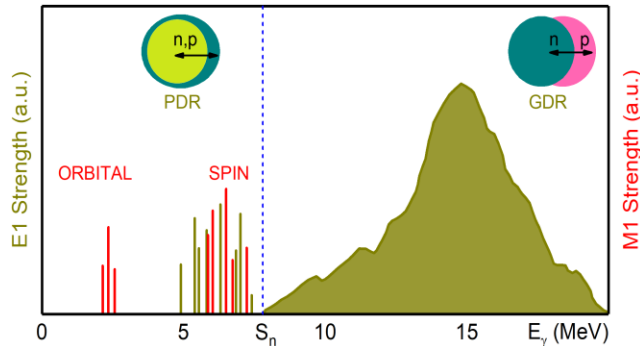
The transition strengths can be extracted in a model-independent way due to the well-known electromagnetic force

$$\sigma_{\mu J}^{(pl)}(E_\gamma) = \frac{8\pi^3 e^2}{3} \frac{(2J+1)(J+1)}{J((2J+1)!!)^2} \left(\frac{E_\gamma}{\hbar c}\right)^{2J-1} \boxed{S_{\mu J}(E_\gamma)}$$

cross section

transition strengths

- **Main modes: Dipole**



$$T_E(L+1)/T_E(L) \sim 10^{-3}$$

$$T_M(L+1)/T_M(L) \sim 10^{-3}$$

$$T_M(L)/T_E(L) \sim 10^{-3}$$

T: transition probability

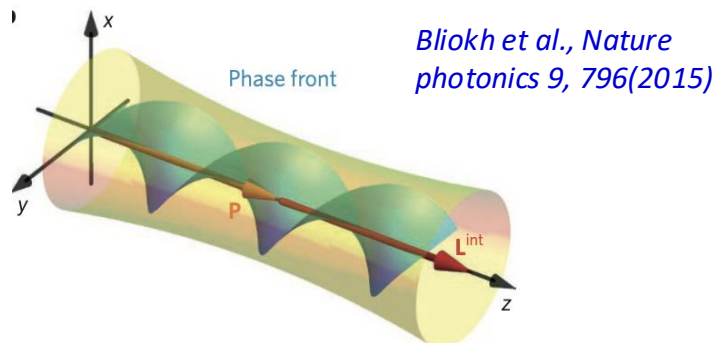
E: electric

M: magnetic

Is it possible to study giant resonances with higher multipolarity by photons?

Vortex gamma photon (with Orbital Angular Momentum)

- Coordinate space**



- ✓ **Vortex wavefunction:**

$$\psi(\mathbf{r}, t) = u(\rho, z) e^{il\phi} e^{ik_z z} e^{-i\omega t}$$

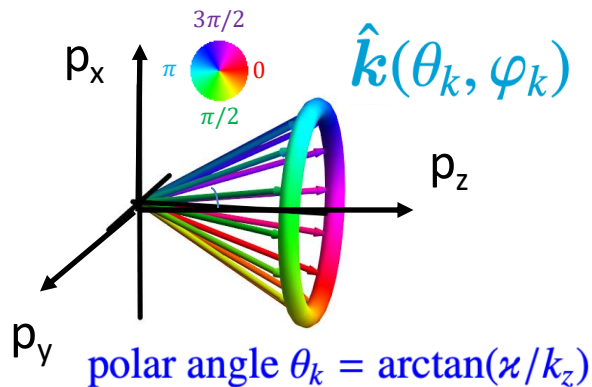
Eigenfunctions for **OAM** operator

$$L_z = -\frac{i\partial}{\partial\phi} \quad \text{OAM } m_l = l\hbar, m_\gamma = m_l + m_s$$

Mode function: $u(\rho, z)$
(Bessel or LG beam modes)

Helical phase: $e^{il\phi} e^{ik_z z}$

- Momentum space**



- ✓ **Superposition of plane waves**

$$\begin{aligned} \psi_{\kappa m k_z}(\mathbf{r}) &= \exp(ik_z z) \int a_{\kappa m}(\mathbf{k}_\perp) \exp(i\mathbf{k}_\perp \rho) \frac{d^2 k_\perp}{(2\pi)^2} \\ &= \int a_{\kappa m}(\mathbf{k}_\perp) \exp(ikr) \frac{d^2 k_\perp}{(2\pi)^2}, \end{aligned}$$

where the Fourier amplitude

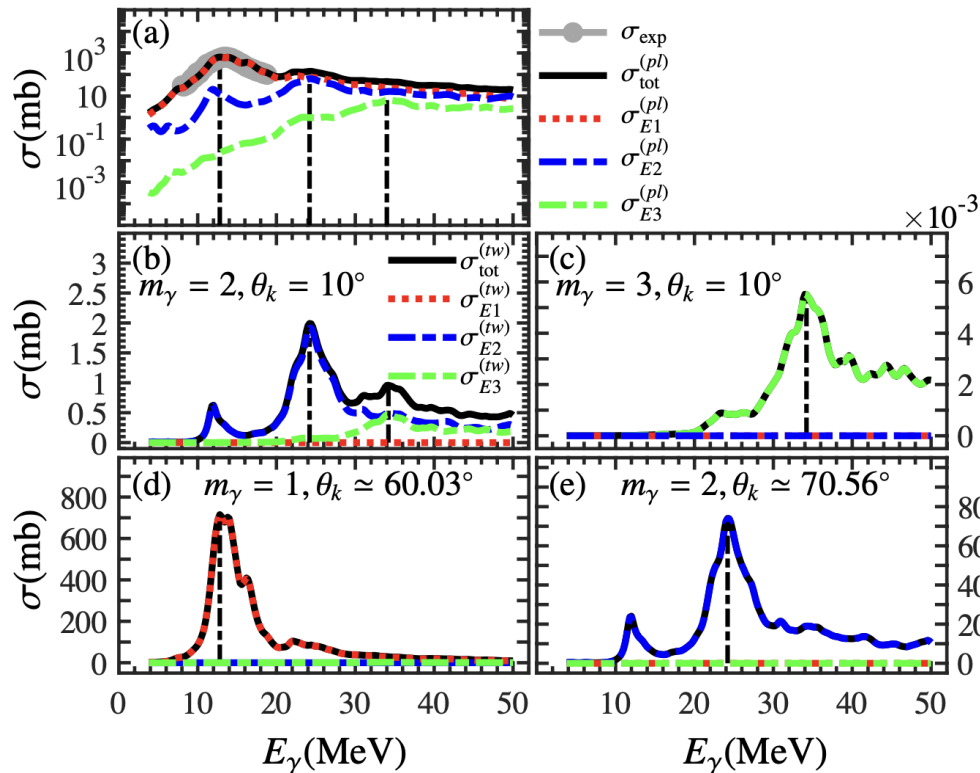
$$a_{\kappa m}(\mathbf{k}_\perp) = i^{-m} \exp(im\varphi_k) \frac{2\pi}{k_\perp} \delta(k_\perp - \kappa)$$

Knyazev and Serbo, Physics uspekhi 61, 449 (2018)

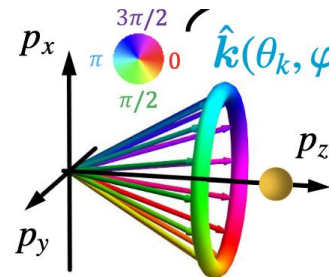
Photo-absorption cross section: plane wave vs. vortex

$\Lambda = 1, \quad b = 0$ RPA+PVC calculation SAMi-T

◆ Plane-wave photon case $M_f - M_i = \Lambda$



◆ Vortex photon case



□ Forbidden of $J < |m_\gamma|$ due to selection rule:

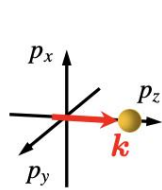
$$M_f - M_i = m_\gamma$$

□ Forbidden of $J = |m_\gamma| + 1$ due to properties of Wigner d-function $d_{m_\gamma \Lambda}^{J_f}(\theta_k)$ which vanishes at specific angle.

Manipulation of giant resonances via vortex photon

Plane wave γ photons:

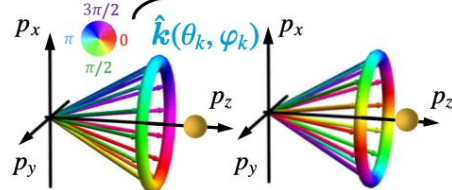
$$M_f - M_i = \Lambda$$



$$m_\gamma = m_s = 1$$

$$m_l = 0$$

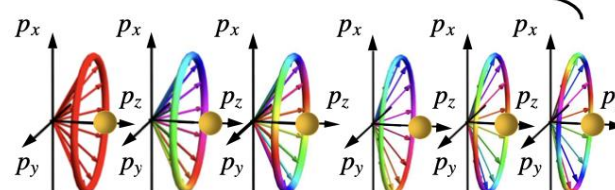
Vortex γ photons: $M_f - M_i = m_\gamma$



small θ_k (as an example)

$$m_\gamma = 2, m_s = m_l \simeq 1$$

$$m_\gamma = 3, m_s \simeq 1, m_l \simeq 2$$

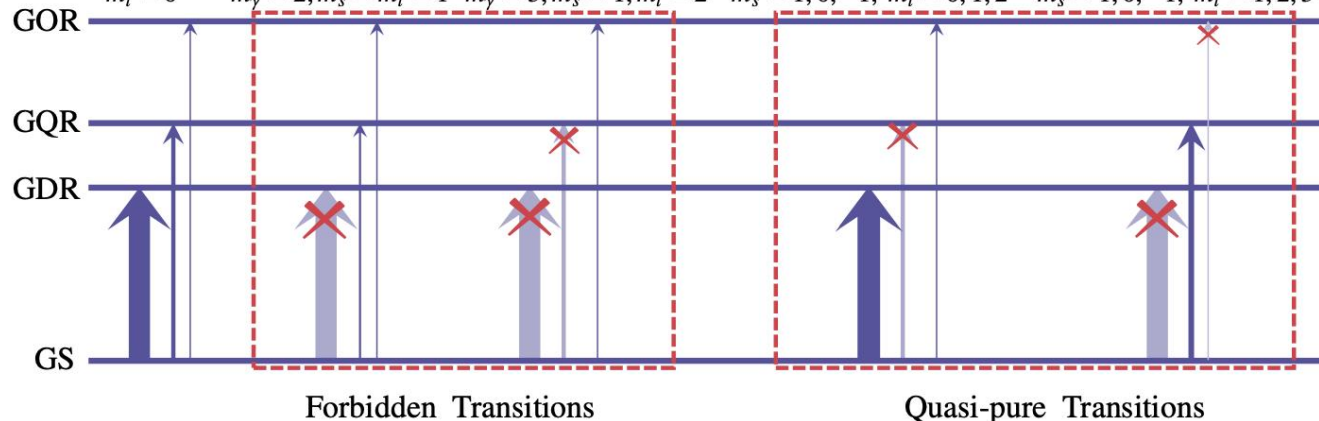


$$m_\gamma = 1, \text{ specific large } \theta_k$$

$$m_s = 1, 0, -1, m_l = 0, 1, 2$$

$$m_\gamma = 2, \text{ specific large } \theta_k$$

$$m_s = 1, 0, -1, m_l = 1, 2, 3$$



Z. W. Lu, L. Guo, Z. Z. Li, M. Ababekri, F. Q. Chen, C. B. Fu, C. Lv, R. R. Xu, X. J. Kong, Y. F. Niu, and J. X. Li, PRL 131, 202502 (2023)

Summary and Perspectives

Summary

- Fully self-consistent QRPA+QPVC based on Skyrme density functional for charge-exchange channel and non-charge-exchange channel are both developed
- ✓ Improve the beta-decay half-lives, which provides accurate inputs for r-process simulation
- ✓ Achieve unified description of GMR in Sn and Pb, which solves long-standing puzzle “Why is Sn so soft?”
- New possibilities to study nuclear collective vibrations with vortex photons are explored.

Perspectives

- Nuclear double beta decay
- Electron-capture in stellar environment
- Neutrino-nucleus scattering

Acknowledgement

Collaborators:

LZU: Z. Z. Li L. Guo X. L. Zhi F. Q. Chen PKU: J. Meng Anhui Uni.: Z. M. Niu

Aizu Univ. and RIKEN: H. Sagawa Milan Univ. : G. Colo, E. Vigezzi

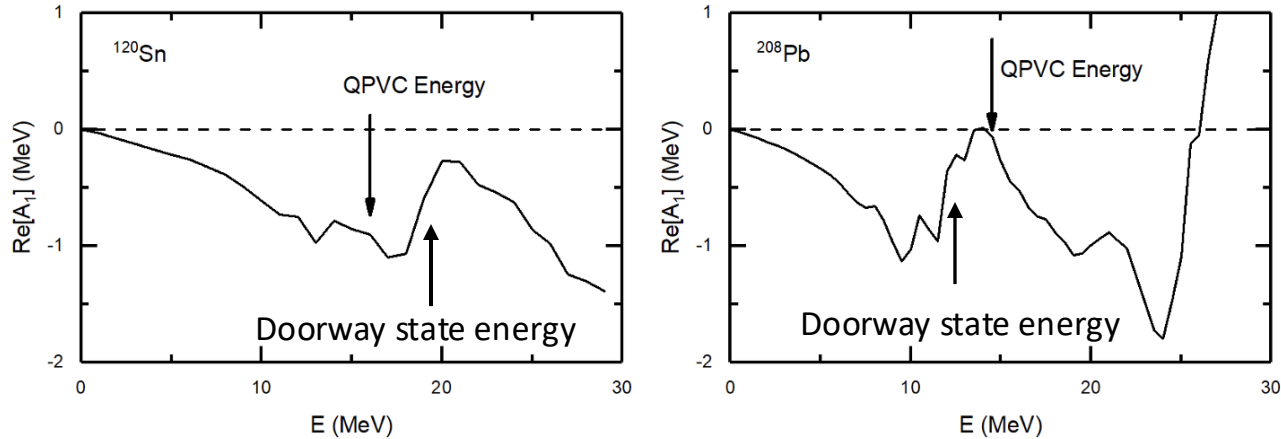
Xian Jiaotong Uni.: Z. W. Lu, M. Ababekri, J. X. Li

CIAE: R. R. Xu, C. Lv Fudan Uni.: C. B. Fu, X. J. Kong

Thank you!

Mechanism: Self-energy

- Real part of self-energy for ^{120}Sn and ^{208}Pb : determines the energy shift

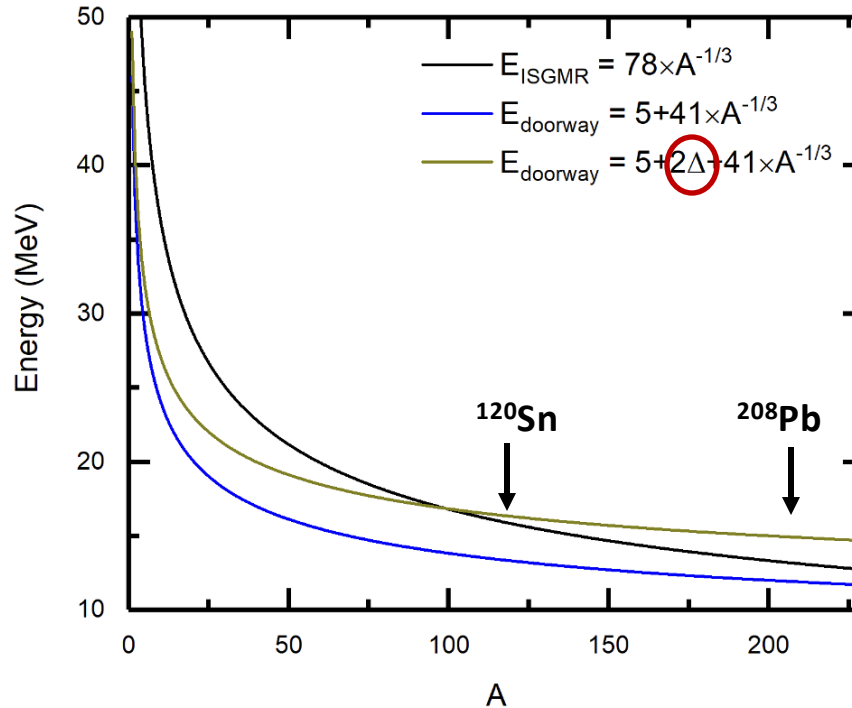


$$\text{Self-energy: } \Sigma(E) \approx \int dE' \frac{V^2}{E - E' + i\epsilon}$$

- E : QPVC energy of GMR peak
- E' : Doorway-state energy (the energy of important doorway state $2\text{qp} \otimes \text{phonon}$)

^{120}Sn : QPVC energy < doorway state energy $\rightarrow$ larger self-energy
 ^{208}Pb : QPVC energy > doorway state energy $\rightarrow$ smaller self-energy

Mechanism: Role of pairing gap



The pairing gap makes the relative energy position of GMR and doorway state different!

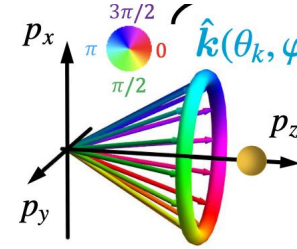
Cross section for vortex photon

- Nuclear photo-absorption cross section for vortex gamma photon

$$\sigma^{(tw)} = \frac{2\pi\delta(E_\gamma - E_f + E_i)}{(2J_i + 1)\bar{J}_z^{(tw)}} \sum_{M_i M_f} M_{M_i M_f}^{(tw)}(\mathbf{b}) M_{M_i M_f}^{(tw)*}(\mathbf{b}).$$

- ✓ Transition amplitude

$$\begin{aligned} M_{M_i M_f}^{(tw)} &= -\frac{1}{c} \langle J_f M_f | \int \hat{\mathbf{j}}(\mathbf{r}) \cdot \mathbf{A}_{\kappa m_\gamma k_z \lambda}^{(tw)}(\mathbf{r}, t) d\mathbf{r} | J_i M_i \rangle \\ &= \int \frac{d^2 k_\perp}{(2\pi)^2} \alpha_{\kappa m_\gamma}(\mathbf{k}_\perp) e^{-i\mathbf{k}_\perp \cdot \mathbf{b}} M_{M_i M_f}^{(pl)}(\theta_k, \varphi_k), \end{aligned}$$



By rotating the nucleus from the propagation axis to the $\mathbf{k}$ direction
(plane-wave component)

$$M_{M_i M_f}^{(pl)}(\theta_k, \varphi_k) = e^{-i(M_f - M_i)\varphi_k} \sum_{M'_i M'_f} d_{M_i M'_i}^{J_i}(\theta_k) d_{M'_i M'_f}^{J_f}(\theta_k) M_{M'_i M'_f}^{(pl)}(0).$$

Finally

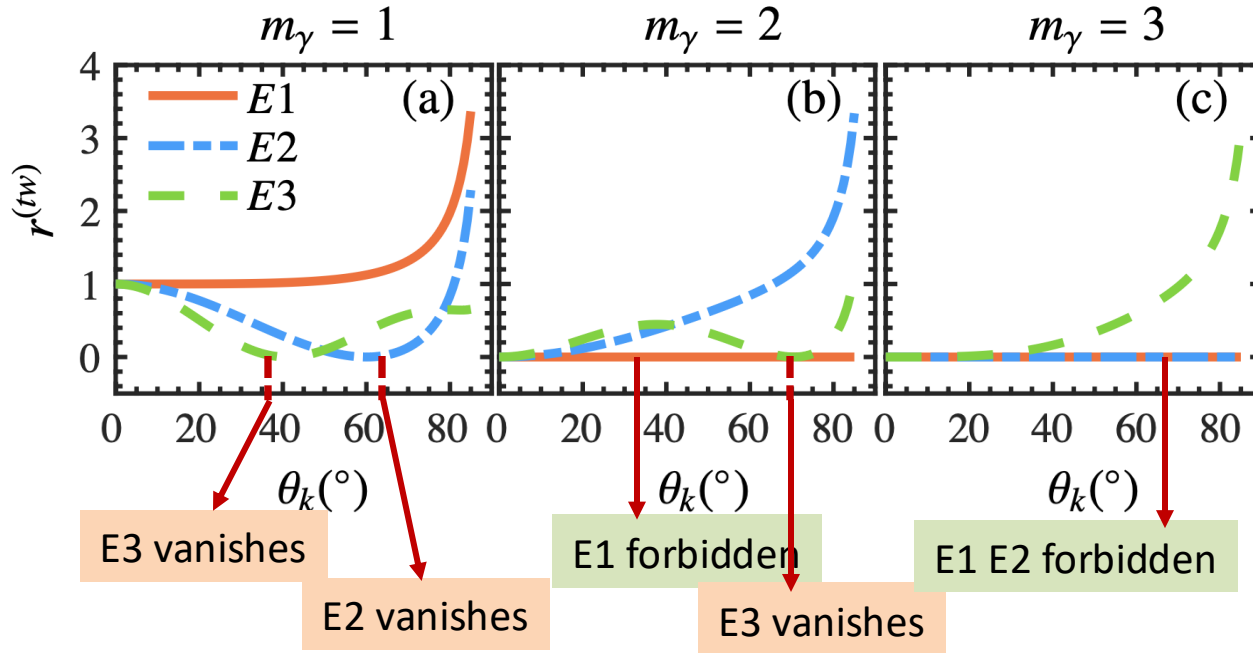
$$M_{M_i M_f}^{(tw)}(\mathbf{b}) = -i^{M_f - M_i - 2m_\gamma} e^{i(m_\gamma + M_i - M_f)\varphi_b} J_{m_\gamma + M_i - M_f}(\kappa b) \sum_{M'_i M'_f} d_{M_i M'_i}^{J_i}(\theta_k) d_{M'_i M'_f}^{J_f}(\theta_k) M_{M'_i M'_f}^{(pl)}(0).$$

- ✓ Average flux density of vortex gamma beam: $\bar{J}_z^{(tw)} = k \cos \theta_k / (2\pi)$

Ratio of the vortex and plane-wave cross section

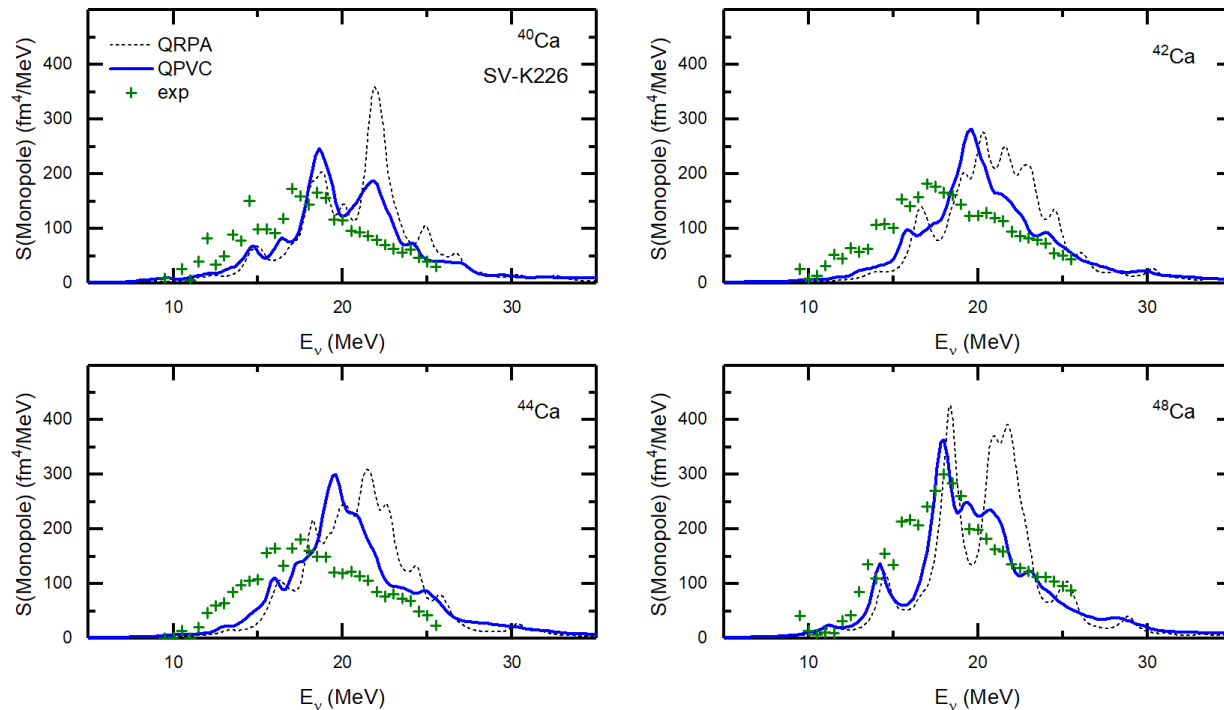
$$\Lambda = 1, \quad b = 0$$

$$\text{polar angle } \theta_k = \arctan(\kappa/k_z)$$



- $m_\gamma = 1, \theta_k = 0$: back to the plane-wave case $r^{(tw)}=1$
- dependence on θ_k comes from Wigner d-function $d_{m_\gamma \Lambda}^{J_f}(\theta_k)$

GMR of Ca isotopes studied by QRPA+QPVC

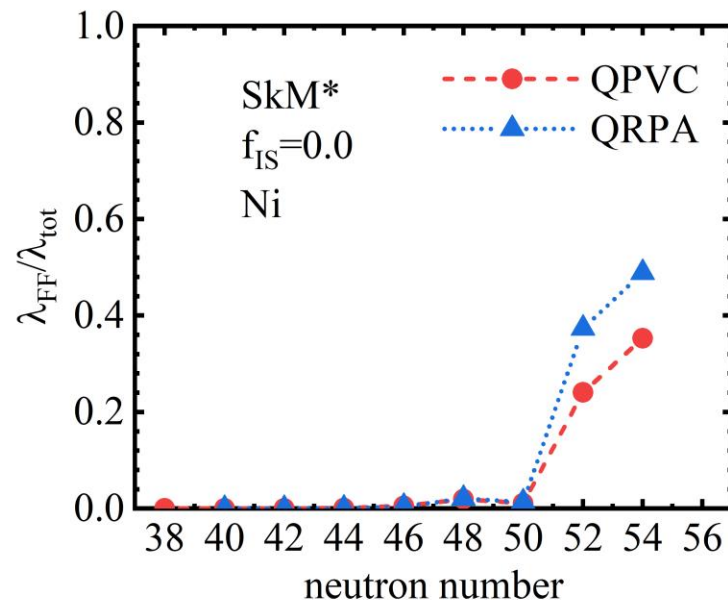
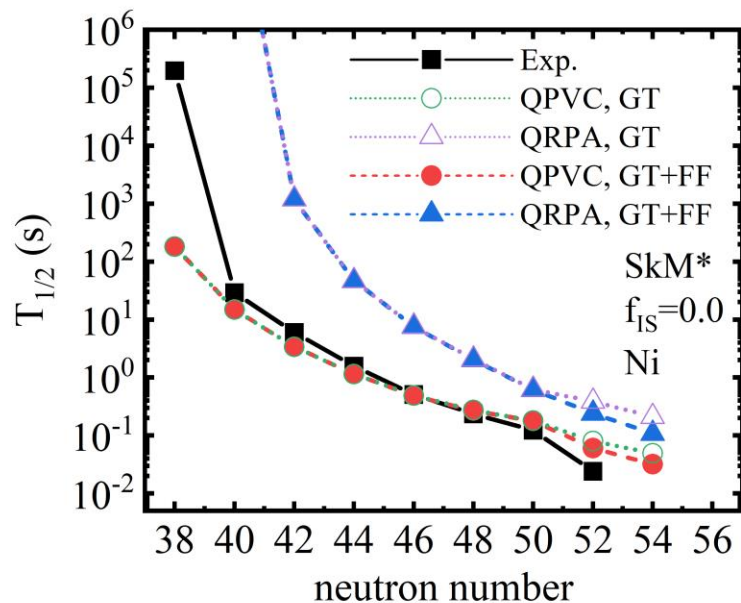


- QRPA:
 - ✓ Much overestimate GMR energies
- QPVC:
 - ✓ Much improve the strength function

Exp.: [Olorunfunmi et al., PRC 105, 054319 \(2022\)](#)

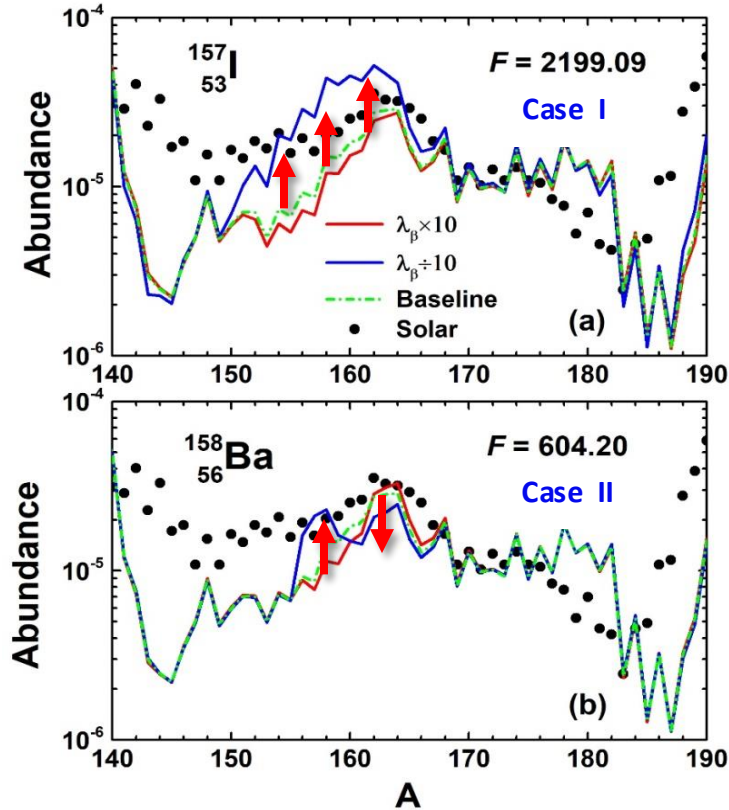
The role of Forbidden Transitions

- Allowed transitions vs. Forbidden transitions
- Contribution of forbidden transitions



- Forbidden transition starts to contribute after N=50 for both QRPA and QPVC
- In QPVC, the contributions from forbidden transitions become smaller

Impact of β -decay half-life on rare-earth peak



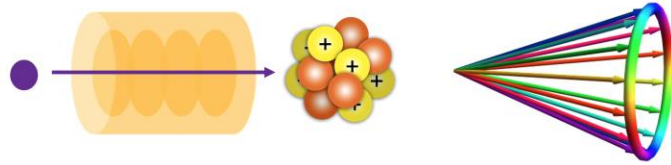
When the β -decay rate decreases:

- **Case I:** Abundance increases in the range of more than 10 mass numbers
- **Case II:** The abundance increases at mass number A of this nucleus and decreases at higher mass numbers in the rare-earth mass region

Y. W. Hao, Y. F. Niu, and Z. M. Niu, PRC 108, L062802 (2023)

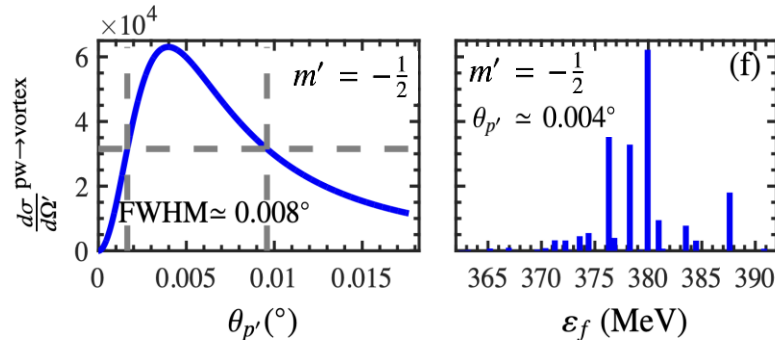
Giant resonances studied by inelastic electron scattering

- Incoming electrons: plane wave (polarized)
- Outgoing electrons: ? Vortex!



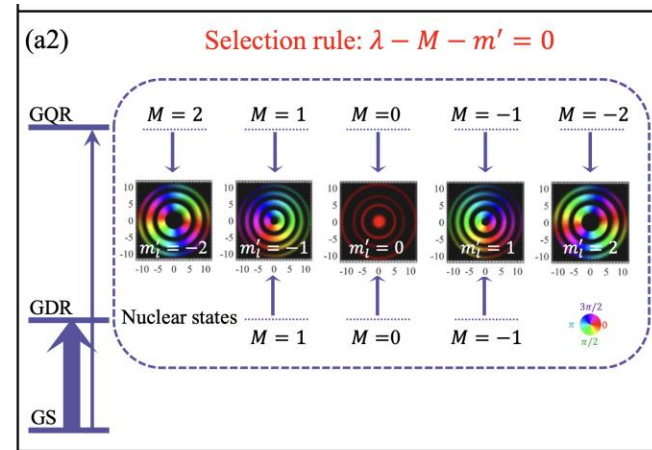
$$\varepsilon_i = 400 \text{ MeV} \quad \lambda = \lambda' = \frac{1}{2}$$

- ✓ Properties of outgoing electrons



- ✓ Selection rule: $\lambda - M - m' = 0$

m' : Angular momentum projection of outgoing electron
 M : Angular momentum projection of giant resonances



By detecting $m' = \lambda \pm 2$, the transition strength of GQR can be extracted model independently.

Z. W. Lu, et al., Phys. Rev. Lett. 134, 052501 (2025)